

Gravitation and quantum interference experiments with neutrons

This article has been downloaded from IOPscience. Please scroll down to see the full text article.

2012 New J. Phys. 14 055010

(<http://iopscience.iop.org/1367-2630/14/5/055010>)

View [the table of contents for this issue](#), or go to the [journal homepage](#) for more

Download details:

IP Address: 64.132.59.114

The article was downloaded on 24/10/2012 at 14:01

Please note that [terms and conditions apply](#).

Gravitation and quantum interference experiments with neutrons

Hartmut Abele¹ and Helmut Leeb

Atominstytut—TU Wien, Stadionallee 2, 1020 Wien, Austria

E-mail: abele@ati.ac.at

New Journal of Physics **14** (2012) 055010 (21pp)

Received 15 August 2011

Published 14 May 2012

Online at <http://www.njp.org/>

doi:10.1088/1367-2630/14/5/055010

Abstract. This paper describes gravity experiments, where the outcome depends upon both the gravitational acceleration g and the Planck constant \hbar . We focus on the work performed with an elementary particle, the neutron.

Contents

1. Introduction	1
1.1. Neutron sources, neutron interferometers and neutron mirrors	2
2. Gravity experiments with neutrons within the classical limit	3
3. Gravitation and neutron-interferometry	4
4. The quantum bouncing ball	7
4.1. Observation of the bound quantum states of neutrons in the gravitational field of the Earth	12
5. The gravity resonance spectroscopy method	14
5.1. Experimental results	17
6. Summary and outlook	19
Acknowledgments	19
References	20

1. Introduction

This paper describes gravity tests with neutrons. We start with a short note on the first free-fall experiment with neutrons in the classical regime. Next we consider experiments where the outcome depends upon both the gravitational acceleration g and the Planck constant \hbar .

¹ Author to whom any correspondence should be addressed.

In 1975, Colella, Overhauser and Werner (COW) [1] demonstrated in their pioneering work a gravitationally induced phase shift. This neutron interferometer experiment was performed at the 2 MW University of Michigan Reactor, and the signal is based on the interference between coherently split and separated neutron de Broglie waves in the gravity potential.

Another idea is to explore a unique system consisting of a single particle, a neutron falling in the gravity potential of the Earth, and a massive object, a mirror, where the neutron bounces off. The task is to study the dynamics of such a quantum bouncing ball, i.e. the measurement of the time evolution of a coherent superposition of quantum states performing quantum reflections [2–4]. In 2002, an experiment showed that it is possible to populate discrete energy levels in the gravitational field of the Earth with neutrons [5]. Successor experiments are performed by the *qBOUNCE* collaboration and the GRANIT collaboration.

The last point is the measurement of the energy splitting between energy eigenstates with a resonance technique [6]. A novelty of this work is the fact that—in contrast to previous resonance methods—the quantum mechanical transition is mechanically driven without any direct coupling of an electromagnetic charge or moment to an electromagnetic potential. Transitions between gravitational quantum states are observed when a Schrödinger wave packet of an ultracold neutron (UCN) couples to the modulation of a hard surface as the driving force. Such experiments operate on the energy scale of pico-eVs and can usefully be employed in measurements of fundamental constants [7] and in a search for non-Newtonian gravity [8].

Gravity tests with neutrons as quantum objects or within the classical limit are reviewed in [9] and our work is an update. The large field of neutron optics and neutron interferometry has been omitted in our work, because it can be found in [10]. Neutron interferometry experiments are also discussed in a review on ‘Neutron interferometry’ [11]. Much of the literature on quantum interference phenomena has been covered in that work.

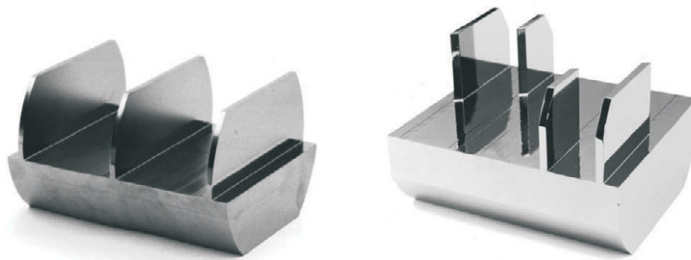
1.1. Neutron sources, neutron interferometers and neutron mirrors

Neutrons are typically produced in a research reactor or a spallation source. Free neutrons cover kinetic energies across the entire energy range from several hundreds of MeV to 10^{20} times less at the pico-eV energy scale. The various fluxes are conveniently defined [12] as fast ($E_n > 1$ MeV), of intermediate energy (1 MeV $\geq E_n > 1$ eV), or slow ($E_n < 1$ eV). Here several subgroups are classified as epithermal (1 eV $\geq E_n > 0.025$ eV), thermal ($\simeq 0.025$ eV), cold (0.025 eV $\geq E_n \geq 0.5$ μ eV), very cold (0.5 μ eV $> E_n \geq 100$ neV) and ultracold (UCN) (< 100 neV) [12] (table 1). From the experimental point of view, ideas have been developed for techniques that would produce a 1000-fold increase in the UCN counting rate due to a higher UCN phase space density. New UCN sources are therefore planned or under construction at many sites, including facilities at FRM-II, ILL, KEK, LANL, Mainz, North Carolina State, PNPI, PSI and Vienna.

Neutrons propagate in condensed matter in a similar manner to the propagation of light, but with a neutron refractive index of less than unity. Interferometers are based on the atom arrangement in silicon crystals and beams are split coherently in ordinary or momentum space for interferometry experiments, where the relative phase becomes a measurable quantity. In this case, nuclear, magnetic and gravitational phase shifts can be applied and measured precisely. A monolithic design and a stable environment provide the parallelity of the lattice planes throughout the interferometer. Figure 1 shows interferometers used in gravity experiments, a symmetric and a skew-symmetric one in order to separate and identify systematic discrepancies

Table 1. From hot to ultracold: neutron energy, temperature, velocity and wavelength distributions.

	Fission neutrons	Thermal neutrons	Cold neutrons	Ultracold neutrons	Gravity experiment
Energy	2 MeV	25 meV	3 meV	<100 neV	1.4 peV (E_{\perp})
Temperature	10^{10} K	300 K	40 K	~ 1 mK	–
Velocity	10^7 m s $^{-1}$	2200 m s $^{-1}$	800 m s $^{-1}$	~ 5 m s $^{-1}$	$v_{\perp} \sim 2$ cm s $^{-1}$
Wavelength		0.18 nm	0.5 nm	~ 80 nm	

**Figure 1.** Photograph of the symmetric (left) and skew-symmetric (right) interferometers used in [22]. The blades of the symmetric interferometer are 3.077 mm thick and 50.404 mm apart. The dimensions of the skew-symmetric interferometer are 16.172 and 49.449 mm. The blade thickness is 2.621 mm. The photograph is reprinted with permission from Littrell K C *et al* 1997 *Phys. Rev. A* **56** 1767 [22]. Copyright 1997 by the American Physical Society.

that depend on the interferometer or the mounting. For each wavelength, the phase difference $\Delta(\lambda)$ was obtained using a technique called the phase rotator interferogram technique. A phase rotator is a 2 mm thick aluminum phase flag, which is placed across both beams and rotated.

The surface of matter constitutes a potential step of height V . Neutrons with transversal energy $E_{\perp} < V$ will be totally reflected. Following a pragmatic definition, UCNs are neutrons that, in contrast to faster neutrons, are retroreflected from surfaces at all angles of incidence. When the surface roughness of the mirror is small enough, the UCN reflection is specular.

This feature makes it possible to build simple and efficient retroreflectors. Such a neutron mirror makes use of the strong interaction between nuclei and UCN, resulting in an effective repulsive force. The potential may also be based on the gradient of the magnetic dipole interaction. Retroreflectors for atoms have used the electric dipole force in an evanescent light wave [13, 14] or are based on the gradient of the magnetic dipole interaction, which has the advantage of not requiring a laser [15].

2. Gravity experiments with neutrons within the classical limit

In the Earth's gravitational field, neutrons fall with an acceleration equal to the local value g [16]. The free fall does not depend on the sign of the neutron's vertical spin component [17]. The studies provide evidence in support of the weak equivalence principle of the equality of inertial mass m_i and gravitational mass m_g . The results obtained by Koester and others confirm that

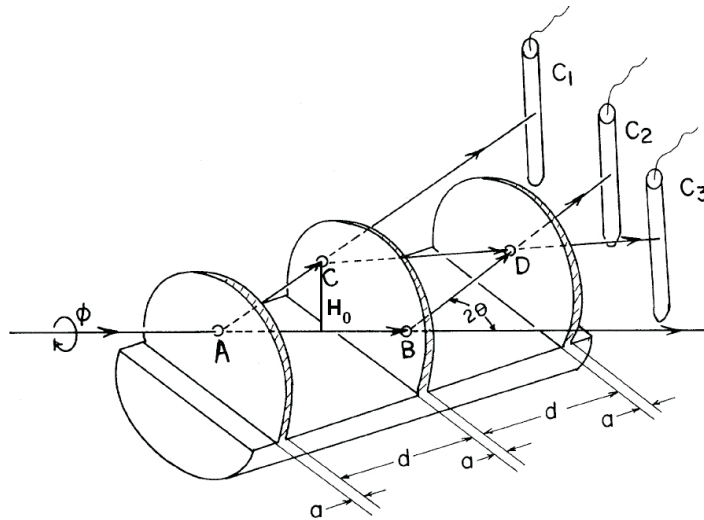


Figure 2. Sketch of the interferometer used in the COW experiment. Reprinted with permission from Collela R, Overhauser A W and Werner S A 1975 *Phys. Rev. Lett.* **34** 1472 [1]. Copyright 1975 by the American Physical Society.

m_i/m_g is equal to unity to an accuracy of 3×10^{-4} [18, 19] and is a consequence of classical mechanics and this has been demonstrated by verifying that neutrons fall parabolically on trajectories in the Earth's gravitational field.

3. Gravitation and neutron-interferometry

COW observed a 'quantum-mechanical phase shift of neutrons caused by their interaction with the Earth's gravitational field' [1]. The signal is based on the interference between coherently split and separated neutron de Broglie waves in the gravity potential. The amplitudes are divided by dynamical Bragg diffraction from perfect silicon crystals. Such interferometers were originally developed for x-rays and then adapted for thermal neutrons by Rauch *et al* [20] in 1974. This neutron interferometer experiment was performed at the 2 MW University of Michigan Reactor.

The following description of the COW experiment, shown in figure 2, follows in part [9, 11]. A monochromatic neutron beam with wavelength λ enters a standard triple-plate interferometer along a horizontal line with momentum $p_0 = \hbar k_0$. The neutron is divided into two wave packets at point A following sub-beam paths ABD and ACD. The interferometer is turned around the incident beam direction by an angle ϕ maintaining the Bragg condition. The neutron wave packets mix and recombine in the third crystal plate at point D, which has a higher gravitational potential above the Earth than the entry point A by a size $m_n g H(\phi) = m_n g H_0 \sin(\phi)$, where m_n is the neutron mass and g is the acceleration of the Earth. The sum of the kinetic energy and the potential energy is constant,

$$E_0 = \frac{\hbar^2 k_0^2}{2m_n} = \frac{\hbar^2 k^2}{2m_n} + m_n g H(\phi), \quad (1)$$

but the difference in height implies that the momentum $p = \hbar k$ on path CD is less than the momentum $p_0 = \hbar k_0$ on path AB. Figure 3 shows the difference in the count rate in counters C2 and C3 as a function of angle ϕ .

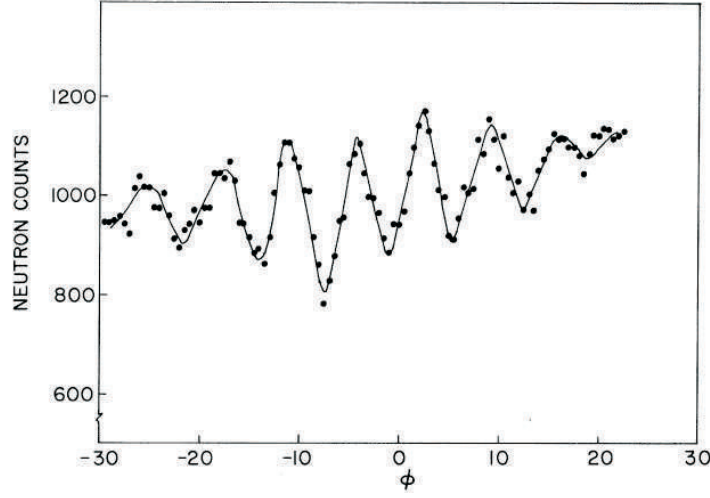


Figure 3. Gravitationally induced quantum interferogram. Reprinted with permission from Collela R, Overhauser A W and Werner S A 1975 *Phys. Rev. Lett.* **34** 1472 [1]. Copyright 1975 by the American Physical Society.

The gravitationally induced quantum interferogram is due to the phase difference of the neutron traversing path ACD relative to path ABD

$$\begin{aligned}\Delta\Phi_{\text{COW}} &= \Phi_{\text{ACD}} - \Phi_{\text{ABD}} \\ &= \Delta k S \\ &\simeq -q_{\text{COW}} \sin \phi,\end{aligned}\quad (2)$$

where $\Delta k = k - k_0$ is the difference in the wave numbers, S the path length of the segments AB and CD and $q_{\text{COW}} = 2\pi\lambda\frac{m_n^2}{\hbar^2}gA_0$. Here, $A_0 = H_0S$ is the area of the parallelogram enclosed by the beam path (figure 2).

The COW measurement differs slightly from the expected value of equation (2) due to additional phase shifts caused by the bending of the interferometer under its weight and by the Earth's rotation, the Sagnac effect. There are small dynamical diffraction effects, too, shifting the central frequency of the oscillations from q_{COW} to

$$q_{\text{grav}} = q_{\text{COW}}(1 + \epsilon), \quad (3)$$

where the correction factor ϵ of several per cent depends on the thickness-to-distance ratio of the interferometer plates. The total phase shift of tilt angle ϕ practically originates from three terms [11],

$$\begin{aligned}\Delta\Phi(\phi) &= \Delta\Phi_{\text{grav}}(\phi) + \Delta\Phi_{\text{bend}}(\phi) + \Delta\Phi_{\text{Sagnac}}(\phi) \\ &= -q_{\text{grav}} \sin \phi - q_{\text{bend}} \sin \phi + q_{\text{Sagnac}} \cos \phi \\ &= q \sin(\phi - \phi_0),\end{aligned}\quad (4)$$

where the frequency of the oscillation with regard to the tilt angle is given by

$$q = [(q_{\text{grav}} + q_{\text{bend}})^2 + (q_{\text{Sagnac}})^2]^{1/2} \quad (5)$$

and

$$\phi_0 = \tan^{-1} \left[\frac{q_{\text{Sagnac}}}{q_{\text{grav}} + q_{\text{bend}}} \right]. \quad (6)$$

Table 2. Eigenenergies and classical turning points for neutrons, atoms and electrons.

	Neutron	^4He	^{85}Ru	^{133}Ce	Electron
E_1 (peV)	1.4	2.3	6.2	7.2	0.12
E_2 (peV)	2.5	3.9	11.0	12.7	0.20
z_1 (μm)	13.7	5.5	0.7	0.5	2061
z_2 (μm)	24.0	9.5	1.2	0.9	3604

A comparison of gravitationally induced quantum interference experiments is given in table 2, providing information on the interferometer parameters, the neutron wavelength λ and the interferogram oscillations q_{COW} . In recent years, these systematic effects have been studied carefully. In the experiment of [21], the bending was separately measured with x-rays and equation (2) reads numerically

$$\begin{aligned}
 q_{\text{grav}} &= (q_{\text{exp}}^2 - q_{\text{Sagnac}}^2)^{1/2} - q_{\text{bend}} \\
 &= (60.12^2 - 1.45^2)^{1/2} - 1.42 \text{ rad} \\
 &= 58.72 \pm 0.03 \text{ rad}.
 \end{aligned}
 \tag{7}$$

The theoretical prediction $q_{\text{grav}} = 59.2 \pm 0.1$ rad is 0.8% higher.

The gravitationally induced phase shift is proportional to λ , and the phase shift due to bending is proportional to λ^{-1} . Therefore, a simultaneous measurement with two neutron wavelengths can be used to determine both contributions.

In the experiment of Littrell *et al* [22], nearly harmonic pairs of neutron wavelengths were used to measure and compensate for effects due to the distortion of the interferometer as it was tilted about the incident beam direction. Figure 1 shows the interferometers used in this experiment, a symmetric and a skew-symmetric one, in order to separate and identify systematic discrepancies that depend on the interferometer or the mounting. For each wavelength, the phase difference $\Delta(\lambda, \phi)$ was obtained using a phase rotator, which is placed across both beams and rotated. A series of phase rotator scans was taken for various values of ϕ using the wavelengths 0.21440 and 0.10780 nm. The corresponding Bragg angles Θ_{B} are 34.15° and 33.94° . The Si(220) or Si(440) Bragg reflection was used. The phase advances by almost the same amount with each step and nearly twice as much for the long wavelength as for the shorter wavelength. Previously, the neutron Sagnac phase shift due to the Earth's rotation was measured [23–25] and was shown to be in agreement with the theory of the order of a few per cent. This phase shift is small for the tilt angles spanned in this experiment and has been subtracted from the data. A historical summary of gravitationally induced quantum interference experiments [21, 22, 25, 26] can be found in [11] (table 3). The most precise measurements made by Littrell *et al* [22] with perfect silicon crystal interferometers show a discrepancy with the theoretical value of the order of 1%.

The experiment of van der Zouw *et al* [26] used very cold neutrons with a mean wavelength of $\lambda = 10$ nm (a velocity of 40 m s^{-1}) and phase gratings as the beam splitting mechanism. The interferometer has a length of up to several meters. The results are consistent with the theory within the measurement accuracy of 1% [26]. The experiments of [22, 25] were performed at the 10 MW University of Missouri Research Reactor.

Table 3. History of gravity-induced interference experiments with symmetric (sym.) and skew-symmetric silicon interferometers. The restricted (rest.) range data mean that the tilt angle $|\phi| = 11^\circ$. The two wavelengths of [22] are diffracted by the (220) or (440) lattice planes. The table is based on [11].

	Interferometer	λ (nm)	A_0 (cm ²)	q_{cow} (theory) (rad)	q_{cow} (exp.) (rad)	Agreement with theory (%)
Collela <i>et al</i> [1]	Sym. number 1	1.445(2)	10.52(2)	59.8(1)	54.3(2.0)	12
Staudenmann <i>et al</i> [25]	Sym. number 2	1.419(2)	10152(4)	56.7(1)	54.2(1)	4.4
		1.060(2)	7,332(4)	30.6(1)	28.4(1)	7.3
Werner <i>et al</i> [21]	Sym. number 2	1.417(1)	10.132(4)	56.50(5)	56.03(3)	0.8
	(440) Full range	1.078(6)	12.016(3)	50.97(5)	49.45(5)	3.0
	Rest. range	1.078(6)	12.016(3)	50.97(5)	50.18(5)	1.5
(220)	Full range	2.1440(4)	11.921(3)	100.57(10)	97.58(10)	3.0
	Rest. range	2.1440(4)	11.921(3)	100.57(10)	99.02(10)	1.5
Littrell <i>et al</i> [22]	Large sym.					
	(440) Full range	1.8796(10)	30.26(1)	223.80(10)	223.38(30)	0.6
	(220) Rest. range	1.8796(10)	30.26(1)	223.80(10)	221.85(30)	0.9

The results are remarkable in some respects. First of all, they demonstrate the validity of quantum theory in a gravitational potential on the 1% level. The gravity-induced interference is purely quantum mechanical, because the phase shift is proportional to the wavelength λ and depends explicitly on Planck's constant \hbar , but the interference pattern disappears as $\hbar \rightarrow 0$. The effect depends on $(m_n/\hbar)^2$ and the experiments test the equivalence principle.

It has been noted that the gravity-induced quantum interference has a classical origin, due to the time delay of a classical particle experienced in a gravitational background field, and classical light waves also undergo a phase shift in traversing a gravitational field [27]. Cohen and Mashhoon [28] have already derived this result with the argument that the index of refraction n in the exterior field of a spherically symmetric distribution of matter is only a function of the isotropic radial coordinate of the exterior Schwarzschild geometry, which is asymptotically flat.

In recent years, ultracold atoms have been used to carry out high-precision measurements with atom interferometers. This technique has been reviewed in [29]. The Stanford group [30] measured g with a resolution of $\Delta g/g = 1 \times 10^{-10}$ after 2 days of integration time. The result has been reinterpreted as a precise measurement of the gravitational redshift by the interference of matter waves [31]. The interpretation has been questioned [33, 34]; see also [32].

4. The quantum bouncing ball

Above a mirror, the gravity potential leads to discrete energy levels of a bouncing massive particle. The corresponding quantum mechanical motion of a massive particle in the gravitational field was named the quantum bouncer [3, 35–37]. The discrete energy levels occur due to the combined confinement of the matter waves by the mirror and the gravitational field. For neutrons the lowest discrete states are in the range of several pico-eVs, opening the way to a new technique for gravity experiments and measurements of fundamental properties.

A quantum mechanical description of UCNs of mass m_n moving in the gravitational field above a mirror is essentially a one-dimensional (1D) problem. The corresponding gravitational

potential is usually given in linear form by $m_n g z$, where g is the gravitational acceleration and z the distance above the mirror, respectively. The mirror, frequently made of glass, with its surface at $z = 0$ is represented by a constant potential V_{mirror} for $z \leq 0$. The potential V_{mirror} is essentially real because of the small absorption cross section of glass and is about 100 neV high, which is large compared to the neutron energy E_{\perp} perpendicular to the surface of the mirror. Therefore it is justified to assume that the mirror is a hard boundary for neutrons at $z = 0$.

The time-dependent Schrödinger equation for the neutron quantum bouncer is given by

$$-\frac{\hbar^2}{2m_n} \frac{\partial^2 \Psi}{\partial z^2} + m_n g z \Psi = i\hbar \frac{\partial \Psi}{\partial t} \quad \text{for } z > 0, \quad (8)$$

where the mirror surface is at $z = 0$. Assuming an infinite mirror potential V_{mirror} implies that $\Psi(z, t)$ vanishes at all times at the mirror surface. The solution of equation (8) is given by a superposition of energy eigenstates

$$\Psi(z, t) = \sum C_n e^{-iE_n t/\hbar} \psi_n(z). \quad (9)$$

The coefficients C_n are determined from the initial condition $\Psi(z, 0)$, and the eigenfunctions ψ_n are the solutions of the time-independent Schrödinger equation

$$\left\{ -\frac{\hbar^2}{2m_n} \frac{\partial^2}{\partial z^2} + V(z) \right\} \psi_n = E_n \psi_n, \quad (10)$$

with

$$V(z) = \begin{cases} m_n g z, & z \geq 0, \\ \infty, & z < 0. \end{cases} \quad (11)$$

This equation describes the neutron in the gravitational field above a mirror at rest. It is convenient to scale (10) with the characteristic gravitational length scale [38] of the bouncing neutron

$$z_0 = \left(\frac{\hbar^2}{2m_n^2 g} \right)^{1/3} = 5.87 \mu\text{m} \quad (12)$$

and the corresponding characteristic gravitational energy scale is

$$E_0 = \left(\frac{\hbar^2 m_n g^2}{2} \right)^{1/3} = 0.602 \text{ peV}. \quad (13)$$

Thus equation (10) takes the concise form

$$\psi''(\zeta) - (\zeta - \zeta_E) \psi(\zeta) = 0, \quad (14)$$

where $\zeta = z/z_0$, $\zeta_E = z_E/z_0$ and $z_E = E/mg$.

For a neutron above a vertical mirror the linear gravity potential leads to the following discrete energy eigenstates: the lowest-energy eigenvalues E_n ($n = 1, 2, 3, 4, 5$) are 1.41, 2.46, 3.32, 4.09 and 4.78 peV. The corresponding classical turning points z_n are 13.7, 24.1, 32.5 and 40.1 μm (see table 2). The energy levels together with the neutron density distribution are shown in figure 4. The idea of observing quantum effects in such a gravitational cavity was discussed for neutrons [39] as well as with atoms [38]. The corresponding eigenenergies and associated classical turning points for states with $n = 1$ and 2 of a neutron, an electron and several atoms are compared in table 2.

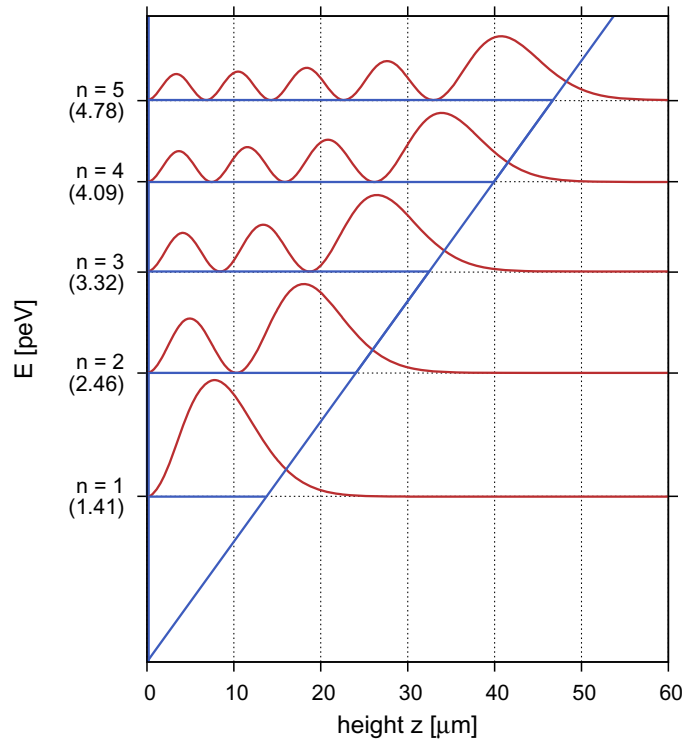


Figure 4. Energy eigenvalues (blue) and neutron density distributions (red) for level 1–5 (figure reproduced with permission from Abele H *et al* 2010 *Phys. Rev. D* **81** 065019 [8]. Copyright 2010 by the American Physical Society).

As Gea-Banacloche [37] has pointed out, the eigenfunctions for this problem are *pieces* of the same Airy function in the sense that they are shifted in each case in order to be zero at $z = 0$ and cut for $z < 0$. Secondly, the wavelength of the oscillations decreases towards the bottom at small z . This is in accordance with the de Broglie relation $\lambda = h/mv$, since the velocity of a classical particle is greater there.

The *qBOUNCE* experiment has been focused on the Quantum Bouncing Ball, i.e. a measurement of the time evolution of the Schrödinger wave function of a neutron bouncing above a mirror [3]. The experiment at the Institut Laue-Langevin, Grenoble, has been performed in the following way [2]: neutrons are taken from the UCN installation PF2. A narrow beam of UCNs is prepared with an adjustable horizontal velocity range between 5 and 12 m s^{-1} . Figure 5 (left) shows a sketch of the experiment. At the entrance of the experiment, a collimator absorber system limits the transversal velocity to an energy in the pico-eV range; see table 2. The experiment itself is mounted on a polished plane granite stone with an active and a passive antivibration table underneath. This stone is leveled using piezo translators. Inclinometers together with the piezo translators in a closed loop circuit guarantee leveling with a precision better than 1 μrad . A solid block with dimensions 10 cm \times 10 cm \times 3 cm composed of optical glass serves as a mirror for neutron reflection. The neutrons see a surface that is essentially flat. An absorber/scatterer, a rough mirror coated with a neutron absorbing alloy, is placed above the first mirror at a certain height in order to select different quantum states. The neutrons are guided through this mirror–absorber system in such a way that they are in the first few quantum states. Neutrons in higher, unwanted states are scattered out of this system. The neutron loss

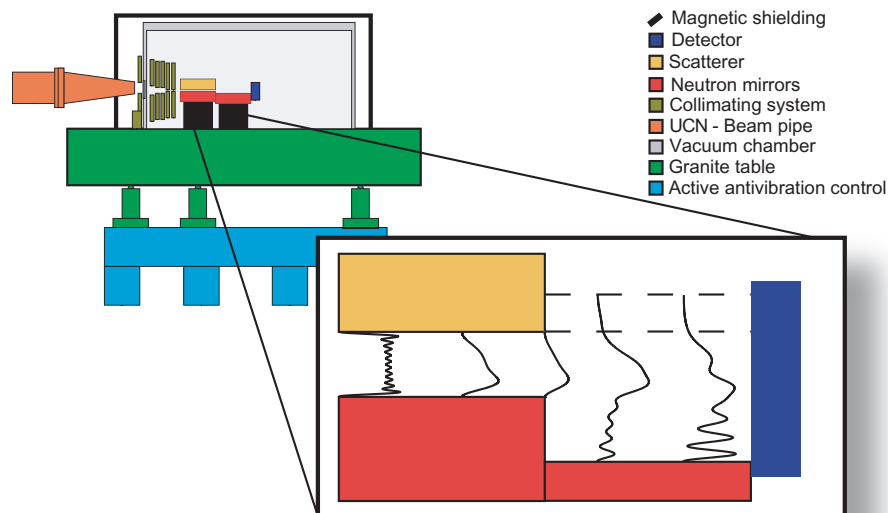


Figure 5. Sketch of the setup. Neutrons are prepared in the lower quantum states and then fall down a step of $27\ \mu\text{m}$ (figure reproduced with permission from Elsevier [2, 3]).

mechanism itself is described at length in [40–42]. This preparation mechanism can be simply modeled by a general phenomenological loss rate $\Gamma_n(l)$ of the n th bound state, which is taken to be proportional to the probability density of the neutrons at the absorber/scatterer.

After this preparation process the quantum state falls a step of several μm . This has been achieved with a second mirror, which is placed after the first one, shifted down by $27\ \mu\text{m}$ (figure 5) [3]. The neutron is now in a coherent superposition of several quantum states. Quantum reflected, the wave function shows aspects of quantum interference, as it evolves with time.

The Schrödinger wave packet—the spatial probability distribution—has been measured at four different positions with a spatial resolution of about $1.5\ \mu\text{m}$. Figure 6 shows a first result with the CR39 detector at positions $x = 0\ \text{cm}$, $x = 6\ \text{cm}$; the quantum fringes are already visible here. The theory makes use of the independently determined parameters for the scatterer and has been convoluted with the spatial detector resolution.

The spatial resolution detector has been developed using ^{10}B -coated organic substrates (CR39) [43]. The detectors operate with an absorptive layer of ^{10}B on plastic CR39. Neutrons are captured in this coated boron layer in an $\text{Li-}\alpha$ -reaction. Boron has two naturally occurring and stable isotopes, ^{11}B (80.1%) and ^{10}B (19.9%). The ^{10}B isotope is ideal in capturing thermal neutrons. The $\text{Li-}\alpha$ -reaction converts a neutron into a detectable track on CR39. An etching technique makes tracks with a length of about $3\ \mu\text{m}$ visible. The information is retrieved optically by an automatic scanning procedure. The microscope is equipped with a micrometer scanning table and a charge coupled device (CCD) camera of sufficient quality. A drawback of this procedure is that it is very time-consuming and does not allow for fast feedback. In addition, corrections are large and of the order of $5\ \mu\text{m}$ to $50\ \mu\text{m}$. An electronic alternative, which allows online access to the data, is modern CCD or CMOS chips. Again, a thin boron layer will serve as a neutron converter.

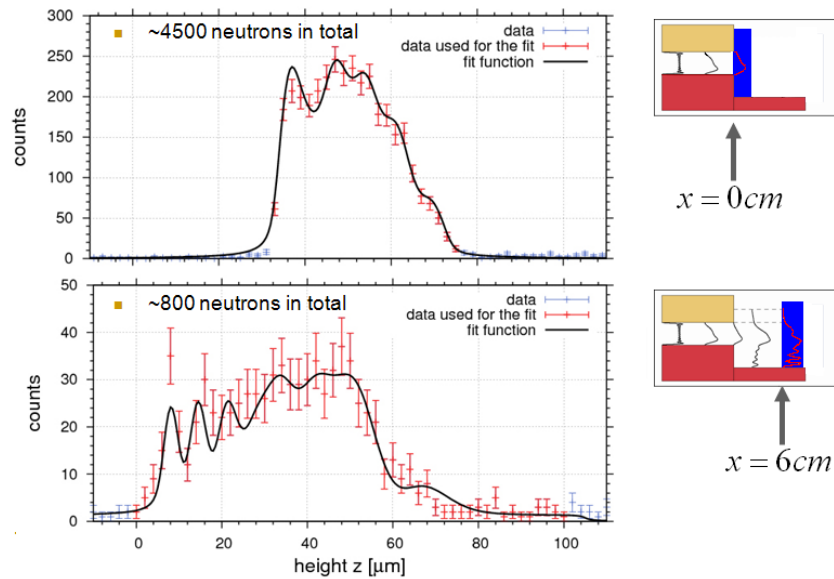


Figure 6. A measurement of the quantum bouncer. Upper figure: a fit to the square of the prepared Schrödinger wave functions just at the step ($x = 0$ cm). Lower figure: quantum prediction after falling and rebounding ($x = 6$ cm). The figure reproduced with permission from Elsevier [3].

A visualization of the neutron motion of a quantum bouncer is very helpful for the interpretation of the experiment. Therefore a simulation [4] of the quantum bouncer has been set up following the theoretical basis outlined in equations (8) and (9). The assumed experimental setup of the simulation is equivalent to that of figure 6, but with a step height of $40 \mu\text{m}$ instead of $27 \mu\text{m}$. At the beginning of the simulation the neutron wave is in the quantum state $n = 1$ above a mirror at height $z = 40 \mu\text{m}$. Assume that a Gaussian wave packet with a width of $1 \mu\text{m}$ and a mean velocity of 6 m s^{-1} incident in the x -direction passes through the edge, after which it has to be represented as a superposition of many discrete eigenstates above the mirror at $z = 0 \mu\text{m}$. In this new representation the largest contributions are from the states with $n = 5, 6, 7$, but contributions greater than 1% can be found up to $n = 19$.

Figure 7 shows the spatial probability distribution with several snapshots of the fall and rise of the wave packet. Just in front of the $40 \mu\text{m}$ step at $t = -0.5$ ms, the Schrödinger wave packet is localized in state $|1\rangle$. At $t = 1.25$ ms the figure shows an intermediate step during the fall; at $t = 3.25$ ms the wave packet is performing a quantum reflection; $t = 5$ ms shows the rise; at $t = 6.5$ ms the turning point is reached; and at $t = 9.75$ ms the wave packet is performing a second reflection.

The neutron quantum bouncer exhibits several interesting quantum phenomena such as collapses and revivals of the wave function. These are a consequence of the different evolutions of the phases of the contributing quantum states. The expectation value of the time evolution shows a pattern similar to quantum beats, but is not harmonic due to the n dependence of energy differences $E_n - E_{n-1}$ in the gravitational potential. Following the evolution in detail, one observes that the first few bounces are clearly visible and follow the classical motion. As time goes on, it is not possible to tell whether the particle is falling down or going up, and the expectation value $\langle z \rangle$ of the wave packet remains very close to the time average of the classical trajectory. Later, the oscillations start again and the particle bounces again. The revival of the

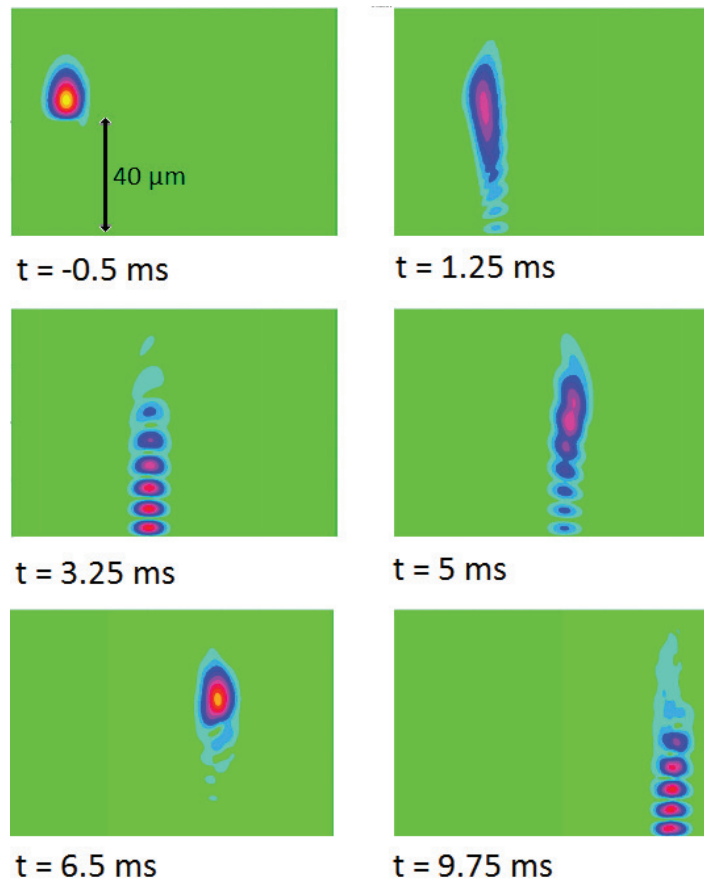


Figure 7. A simulation of the quantum bouncing ball at different time steps. A neutron in quantum state $|1\rangle$ is localized at a height of $40\ \mu\text{m}$ and then falls, reflects and bounces back (figure reproduced with permission from [4]).

oscillation is a purely quantum phenomenon and has no simple quasiclassical explanation. For more details, see [4].

4.1. Observation of the bound quantum states of neutrons in the gravitational field of the Earth

The first observation of quantum states in the gravitational potential of the Earth with UCNs [5, 42, 44, 45] was performed at the Institut Laue-Langevin and started as a collaboration between ILL (Grenoble), PNPI (Gatchina), JINR (Dubna) and the Physics Institute at Heidelberg University. In this experiment, neutrons were allowed to populate the lower states of figure 4. Higher, unwanted states were removed. The neutrons pass through the mirror-absorber/scatterer system shown in figure 8 and are eventually detected by a ^3He counter. The height of the absorber above the mirror is varied and the transmission is measured as a function of height.

The data points, plotted in figures 9 and 10, show the measured number of transmitted neutrons for an absorber height l and follow the expected behavior: no neutrons reach the detector below an absorber height of $15\ \mu\text{m}$. Then, above an absorber height of $15\ \mu\text{m}$, one sees the transmission of ground-state neutrons resulting in an increased count rate. From the classical point of view, the transmission F of neutrons is proportional to the phase space volume allowed

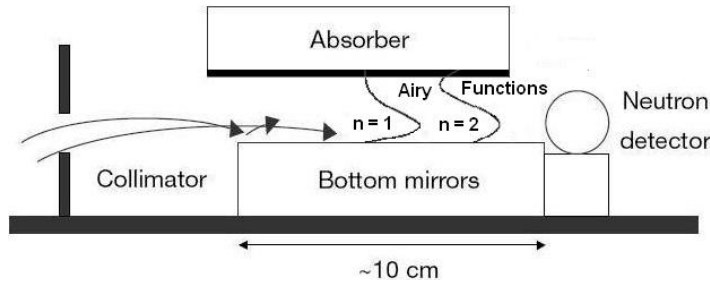


Figure 8. Sketch of the setup of the first observation of bound quantum states in the gravitational field of the Earth. The mirror and the absorber system with adjustable slit height are shown.

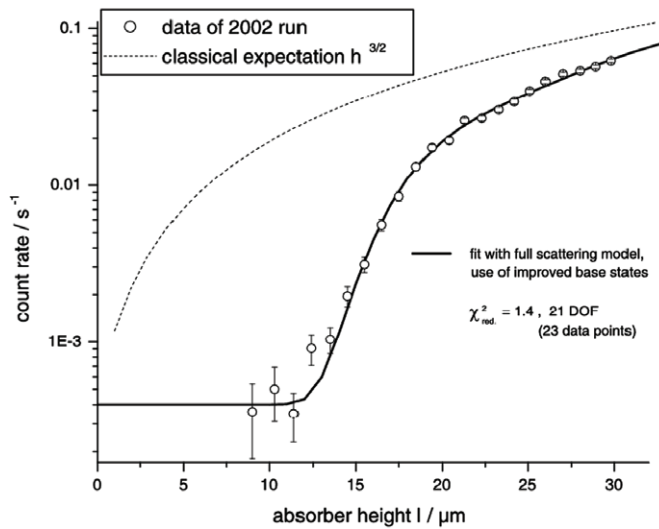


Figure 9. Data 2002 [5] in the logarithmic scale: a fit to the transmission as a function of the absorber height $h = l$ [42]. The difference between the quantum observation and the classical expectation is clearly visible.

by the absorber. It is governed by a power law $F \sim l^n$ and $n = 3/2$. The dotted line shows this classical expectation. The difference between the quantum observation and the classical expectation is clearly visible.

A description of the measured transmission F versus the absorber height l can be obtained via the solution of the decoupled 1D stationary Schrödinger equation in equation (11). The general solution of equation (11) for ψ_n is a superposition of the Airy functions $Ai(z)$ and $Bi(z)$,

$$\psi_n(z) = A_n Ai(\zeta - \zeta_n) + B_n Bi(\zeta - \zeta_n) \quad (15)$$

with the scaling ζ and a displacement ζ_E of equation (14). There is an additional boundary condition at the absorber and we need the term with Airy's Bi -function going to infinity at large z . Due to the absorber, the bound neutron states ψ_n in the linear gravitational potential are confined by two very high potential steps above (absorber) and below (mirror). In the case of $l \leq z_n$, the absorber/scatterer will begin to change the bound states. This implies an energy shift leading to $E_n(l) \gg mg \cdot l$ [40]. Therefore, the calculation of the loss rate is performed using

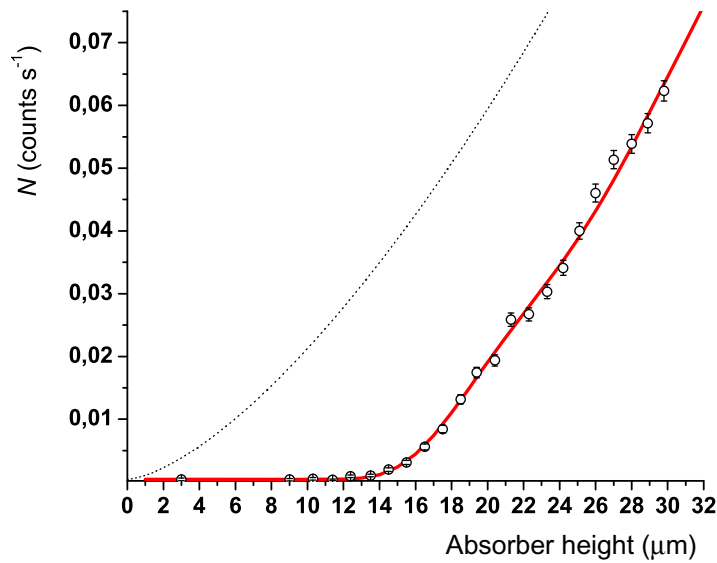


Figure 10. Data 2005 [45] in the linear scale: a fit to the transmission as a function of the absorber height l [42]. The dotted line shows the classical expectation.

the full realistic bound states of equation (15). Only for large l , the n th eigenfunction of the neutron mirror system reduces to the Airy function $\text{Ai}(\zeta - \zeta_n)$ for $n > 0$. The corresponding eigenenergies are $E_n = m_n g z_n$ with $z_n = z_0 \zeta_n$. In the WKB approximation, one obtains in leading order

$$\zeta_n = \left(\frac{3\pi}{2} \left(n - \frac{1}{4} \right) \right)^{2/3}, \quad (16)$$

where z_n corresponds to the turning point of a classical neutron trajectory with energy E_n .

In a different setup, the absorber was placed at the bottom and the mirror on top. Using this reversed geometry, the scattering-induced roughness was found at $z = 0$. The bound states are now, again, given by the Airy functions, but the gravitationally bound states will be strongly absorbed at arbitrary heights l of the mirror at the top—in contrast to the normal setup. Thus, the dominating effect of gravity on the formation of the bound states has been demonstrated, since the measurement did not show just a simple confinement effect [42].

5. The gravity resonance spectroscopy method

A previously developed resonance spectroscopy technique [6] allows a precise measurement [8] of the energy levels of quantum states in the gravity potential of the Earth. Quantum mechanical transitions with a characteristic energy exchange between an externally driven modulator and the energy levels are observed on resonance. An essential novelty of this kind of spectroscopy is the fact that the quantum mechanical transition is mechanically driven by an oscillating mirror and is not a consequence of a direct coupling of an electromagnetic charge or moment to an electromagnetic field. On resonance energy transfer is observed according to the energy difference between gravity quantum states coupled to the modulator. The physics behind these

transitions is related to earlier studies of the energy transfer when matter waves bounce off a vibrating mirror [49, 50] or a time-dependent crystal [51, 52].

The concept is related to Rabi's magnetic resonance technique for the measurements of nuclear magnetic moments [46]. The sensitivity is extremely high, because a quantum mechanical phase shift is converted into a frequency measurement. The sensitivity of resonance methods achieved so far [47] is 6.8×10^{-22} eV and has been achieved in the search for a non-vanishing electric dipole moment of the neutrons. Such an uncertainty in energy corresponds to a time uncertainty of 6 days according to Heisenberg's principle.

In a two-level spin-1/2-system coupled to a resonator, magnetic transitions occur when the oscillator frequency ω is equal to the Bohr frequency of the system. Rabi resonance spectroscopy measures the energy difference between the levels $|p\rangle$ and $|q\rangle$ and damping γ . The wave function of the two-level system is

$$\Psi(\vec{r}, t) = \langle \vec{r} | \Psi(t) \rangle = C_p(t) e^{-i\omega_p t} u_p(\vec{r}) + C_q(t) e^{-i\omega_q t} u_q(\vec{r}) \quad (17)$$

with the time varying coefficients $C_p(t)$ and $C_q(t)$.

With the frequency difference ω_{pq} between the two states, the frequency ω of the driving field, the detuning $\delta\omega = \omega_{pq} - \omega$, the Rabi frequency Ω_R and the time t , the coupling between the time varying coefficients is given by

$$\frac{d}{dt} \begin{pmatrix} \tilde{C}_p(t) \\ \tilde{C}_q(t) \end{pmatrix} = \frac{i}{2} \begin{pmatrix} \delta\omega & \Omega_R \\ \Omega_R^* & -\delta\omega \end{pmatrix} \begin{pmatrix} \tilde{C}_p(t) \\ \tilde{C}_q(t) \end{pmatrix} \quad (18)$$

with a transformation into the rotating frame of reference:

$$\begin{aligned} C_p(t) &= \tilde{C}_p(t) e^{-\frac{i}{2}\delta\omega t}, \\ C_q(t) &= \tilde{C}_q(t) e^{-\frac{i}{2}\delta\omega t}. \end{aligned} \quad (19)$$

Here, Ω_R is a measure of the strength of the coupling between the two levels and is related to the vibration strength. Such oscillations are damped out and the damping rate depends on how strongly the system is coupled to the environment.

In generalizing this system, it is possible to describe quantum states in the gravity field of the Earth in analogy to a spin-1/2-system, where the time development is described by the Bloch equations. Neutron matter waves are excited by an oscillator coupled to quantum states with transitions between the state $|p\rangle$ and the state $|q\rangle$. The energy scale is the pico-eV scale.

In the experiment, the damping is caused by the scatterer [42] at height l above a mirror. A key point for the demonstration of this method is that it allows for the detection of resonant transitions $|p\rangle \rightarrow |q\rangle$ at a frequency which is tuned by this scatterer height. As explained, the additional mirror potential shifts the energy of the state $|3\rangle$ as a function of height; see figure 11. The absorption is described phenomenologically by adding decay terms γ_p and γ_q to the equations of motion:

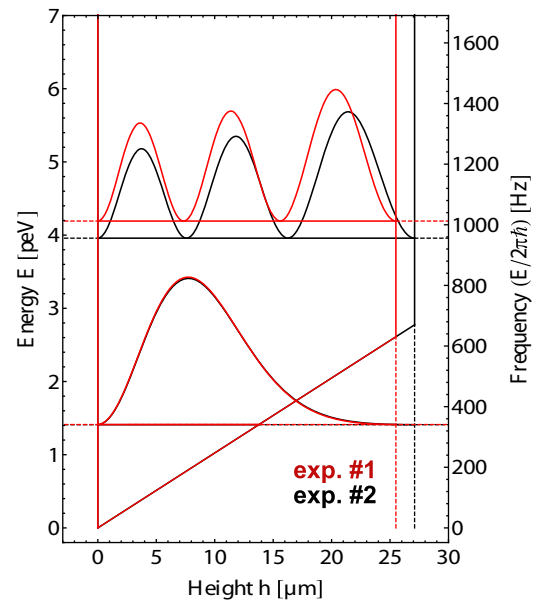
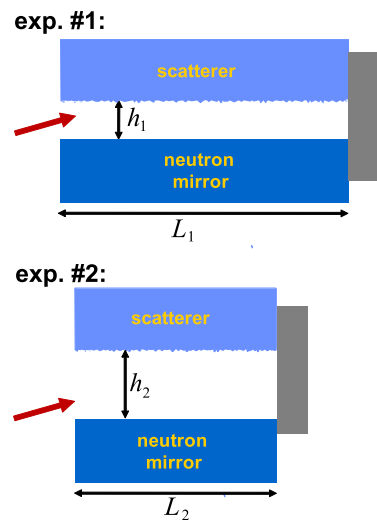
$$\frac{d}{dt} \begin{pmatrix} \tilde{C}_p(t) \\ \tilde{C}_q(t) \end{pmatrix} = \frac{i}{2} \begin{pmatrix} (\delta\omega + i\gamma_p) & \Omega_R \\ \Omega_R^* & (-\delta\omega + i\gamma_q) \end{pmatrix} \begin{pmatrix} \tilde{C}_p(t) \\ \tilde{C}_q(t) \end{pmatrix}. \quad (20)$$

The oscillator is realized by a vibrating mirror, i.e. a modulation of the hard surface potential in vertical position. Neutron mirrors are made of polished optical glass. Interactions limit the lifetime of the state of a two-level system. Lifetime limiting interactions are described phenomenologically by adding decay terms to the equations of motion leading to damped Rabi oscillations. Another concept has been proposed by Kreuz *et al* [48].

Experimental Setup

Height $h \longleftrightarrow$ Energy E

Length $L \longleftrightarrow$ Resonance width



Setup Parameters and Experimental Results

	Length of the neutron mirror	Height of scatterer	Mean time of flight	Energy difference	Resonance frequency (prediction)	Resonance frequency (measurement)	Resonance width (FWHM)
	Length L [cm] Width W [cm] Height H [cm]	h [μm]	t [ms]	E_{13} [peV]	ω_{13} [s ⁻¹]	ω_{13} [s ⁻¹]	$\Delta\omega$ [s ⁻¹]
Experiment 1	15 x 3 x 3	25.5	23	2.78	$2\pi \times 671$	$2\pi \times (705 \pm 6)$	$2\pi \times 41.2$
Experiment 2	10 x 3 x 3	27.1	15	2.55	$2\pi \times 615$	$2\pi \times (592 \pm 11)$	$2\pi \times 61.6$

Figure 11. Experimental parameters (figure reproduced with permission from Nature Physics [6]).

A typical Rabi resonance spectroscopy experiment consists of three regions that the particles pass through. Firstly, one has to install a state selector in region I, secondly, a coupling region inducing specific transitions in the two-state systems, whose energy difference is measured in region II, and a state detector in region III, see figure 12 (left). A so-called π -pulse in region II creates an inversion of the superposition of the two states, whose energy difference is to be measured. Technical details are given in [8].

In this experiment with neutrons, regions I–III are achieved with only one bottom mirror coupled to a mechanical oscillator, a scatterer on top and a neutron detector behind; see figure 12 (right). The scatterer allows only the ground state to pass through and prepares the state $|p\rangle$. It removes and absorbs higher, unwanted states [42] as explained. The vibration, i.e. a modulation of the mirror's vertical position, induces transitions to $|q\rangle$, which are filtered out by the scatterer again. The neutrons are taken from the UCN installation PF2 at Institut Laue-Langevin. The horizontal velocity is restricted to $5.7 \text{ m s}^{-1} < v < 7 \text{ m s}^{-1}$. The experiment itself is mounted on a polished plane granite stone with an active and a passive anti-vibration table underneath. This stone is leveled with a precision better than $1 \mu\text{rad}$. A mu-metal shield suppresses the coupling of residual fluctuations of the magnetic field to the magnetic moment of the neutron sufficiently.

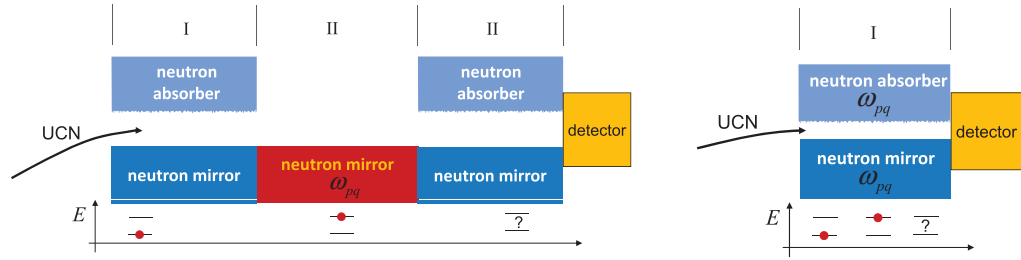


Figure 12. Comparison of two different resonance spectroscopy techniques (figure reproduced with permission from [58]).

5.1. Experimental results

Within the qBounce experiment [53], several resonance spectroscopy measurements with different geometric parameters have been carried out, resulting in different resonance frequencies and widths. In general, the oscillator frequency at resonance for a transition between states with energies E_p and E_q is

$$\omega_{pq} = \frac{E_q - E_p}{\hbar} = \omega_q - \omega_p. \quad (21)$$

The transfer is referred to as the Rabi transition. The transitions $|1\rangle \leftrightarrow |2\rangle$, $|1\rangle \leftrightarrow |3\rangle$, $|2\rangle \leftrightarrow |3\rangle$ and $|2\rangle \leftrightarrow |4\rangle$ have been measured. In detail, the $|1\rangle \leftrightarrow |3\rangle$ transition with $\omega_{13} = \omega_3 - \omega_1$ is described (see the table in figure 11). On resonance ($\omega = \omega_{13}$), this oscillator drives the system into a coherent superposition of states $|1\rangle$ and $|3\rangle$ and one can choose the amplitude a in such a way that we have complete reversal of the state occupation between $|1\rangle$ and $|3\rangle$. It is—as it has been done—convenient to place the scatterer at a certain height h on top of the bottom mirror. This allows us to tune the resonance frequency between $|1\rangle$ and $|3\rangle$ due to the additional potential of the scatterer, which shifts the energy of states $|2\rangle$ and $|3\rangle$, but leaves state $|1\rangle$ unchanged; see figure 11. The energy levels and the probability density distributions for these states are also given in figure 11. The scatterer removes neutrons from the system and the Rabi spectroscopy contains a well-defined damping.

The observable is the measured transmission $|1\rangle$ to $|3\rangle$ as a function of the modulation frequency and amplitude, see figures 13 and 14. For this purpose, piezo-elements have been mounted underneath. They induce a rapid modulation of the surface height with acceleration a , which was measured with a noise and vibration analyzer attached to the neutron mirror system. In addition to this, the position-dependent mirror vibrations were measured using a laser-based vibration analyzer attached to the neutron mirror systems.

For the first experiment, figure 13 shows the measured count rate as a function of ω . Blue (brown) data points correspond to measurements with moderate (high) vibration strength $1.5 \leq a \leq 4.0 \text{ m s}^{-2}$ ($4.9 \leq a \leq 7.7 \text{ m s}^{-2}$). The corresponding Rabi resonance curve was calculated using their mean vibration strength of 2.95 m s^{-2} (5.87 m s^{-2}). The black data point sums up all of the measurements at zero vibration. The gray band represents the one sigma uncertainty of all off-resonant data points. The brown line is the quantum expectation as a function of oscillator frequency ω for Rabi transitions between state $|1\rangle$ and state $|3\rangle$ within an average time of flight $\tau = L/v = 23 \text{ ms}$. The normalization for transmission T , the frequency at resonance ω_{13} and the global parameter f are the only fit parameters. It was found that the vibration amplitude does not change on the flight path of the neutrons but depends in a linear way

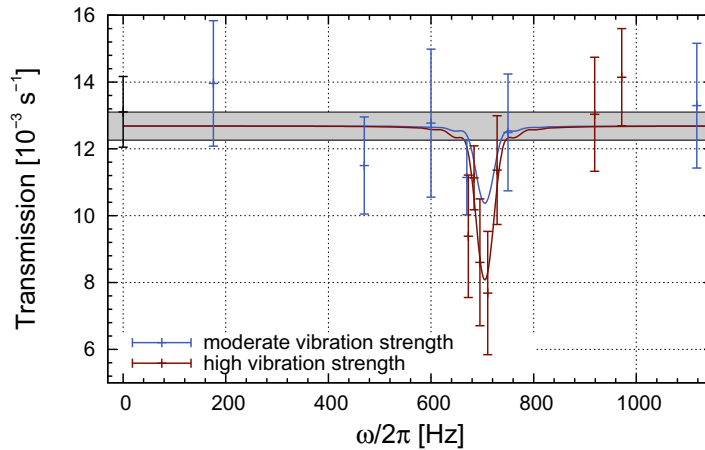


Figure 13. Gravity resonance spectroscopy and excitation [6]: the transmission as a function of modulation frequency shows a sharp resonance at $\omega_{13} = 2\pi \times (705 \pm 6) \text{ s}^{-1}$. The gray band represents the statistical 1σ uncertainty of all off-resonant data points. Blue (brown) data points correspond to measurements with moderate (high) vibration strength $1.5 \leq a \leq 4.0 \text{ m s}^{-2}$ ($4.9 \leq a \leq 7.7 \text{ m s}^{-2}$). The corresponding Rabi resonance curve is calculated using their mean vibration strength of 2.95 m s^{-2} (5.87 m s^{-2}). The black data point sums up all measurements at zero vibration (figure reproduced with permission from Nature Physics [6]).

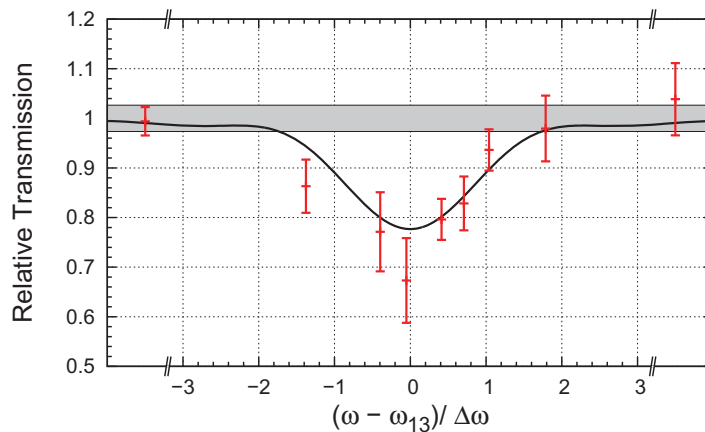


Figure 14. The combined result for both measurements with mirror lengths $L = 10 \text{ cm}$ and $L = 15 \text{ cm}$. The transmission in units of the unperturbed system is displayed as a function of detuning. The significance for gravity spectroscopy between states $|1\rangle$ and $|3\rangle$ at ω_{13} is 4.9 standard deviations. The left and right data points combine all off-resonant measurements with $|(\omega - \omega_{13}) / \Delta\omega| \geq 3$ (figure reproduced with permission from Nature Physics [6]).

on their transversal direction. f is a weighting parameter to be multiplied with the measured vibration strength to correct for these linear effects. A sharp resonance was found at frequency $\omega_{13} = 2\pi \times (705 \pm 6) \text{ Hz}$, which is close to the frequency prediction of $\omega_{13} = 2\pi \times 671 \text{ Hz}$, if we remember that the height measurement has an uncertainty due to the roughness of the

scatterer. For the weighting factor we found $f = 0.56 \pm 0.16$. The full-width at half-maximum is the prediction made from the time the neutrons spend in the modulator. The significance for $|1\rangle \rightarrow |3\rangle$ excitations is 3.5 standard deviations.

The fit used in figure 13 contains three parameters, the resonant frequency ω_{pq} , the transmission normalization N and the vibration strength parameter f to be multiplied with the measured acceleration. The damping as a function of Ω_R was measured separately, see figure 14, as well as the width of the Rabi oscillation and the background ($0.005 \pm 0.0002 \text{ s}^{-1}$).

In a second measurement, the length $L = 10 \text{ cm}$ reduces the average flight time to $\tau = 15 \text{ ms}$. Furthermore, the scatterer height differs by $1.6 \mu\text{m}$ from the first measurement, thus changing the resonant frequency prediction to $\omega_{13} = 2\pi \times 615 \text{ Hz}$. The resonance frequency $\omega_{13} = 2\pi \times (592 \pm 11) \text{ Hz}$ is close to the prediction and $f = 0.99 \pm 0.29$ are observed. Figure 14 shows the combined result for both measurements with mirror lengths $L = 10 \text{ cm}$ and $L = 15 \text{ cm}$. The transmission in units of the unperturbed system is displayed as a function of detuning. In total, the significance for gravity resonance spectroscopy between states $|1\rangle$ and $|3\rangle$ at ω_{13} is 4.9 standard deviations. The left and right data points combine all off-resonant measurements with $|(\omega - \omega_{13}) / \Delta\omega| \geq 3$, where $\Delta\nu$ is the half-width at half-maximum.

6. Summary and outlook

Gravity experiments with neutrons are motivated mainly by the fact that in contrast to atoms the electrical polarizability of neutrons is extremely small. Neutrons are not disturbed by short-range electric forces such as van der Waals or Casimir forces. Together with its neutrality, a neutron provides a sensitivity many orders of magnitude below the strength of electromagnetism.

For the experiments discussed, the preliminary estimated sensitivity of the measured energy difference between the gravity levels is 7.6×10^{-3} . This corresponds to $\Delta E = 2 \times 10^{-14} \text{ eV}$, which tests Newton's law on the micrometer distance at this level of precision.

This accuracy range is of potential interest, because it addresses some of the unresolved questions of science: the nature of the fundamental forces and underlying symmetries and the nature of gravitation at small distances [54]. Hypothetical extra dimensions, curled up to cylinders or tori with a small compactification radius, would lead to deviations from Newton's gravitational law at very small distances [53]. Another example is suggested by the magnitude of the vacuum energy in the universe [55, 56], which again is linked to the modification of gravity at small distances. Furthermore, the experiments have the potential to test the equivalence principle [57].

The long-term plan is to apply Ramsey's method for separated oscillating fields to the spectroscopy of the quantum states in the gravity potential above a horizontal mirror [8]. Such measurements with UCNs will offer a sensitivity to Newton's law or hypothetical short-ranged interactions that is about 21 orders of magnitude below the energy scale of electromagnetism.

Acknowledgments

We acknowledge support from the Austrian Science Fund (FWF) under contract no. I529-N20, I534-N20 and from the German Research Foundation (DFG) within the Priority Programme

(SPP) 1491 ‘Precision experiments in particle and astrophysics with cold and ultracold neutrons’, the DFG Excellence Initiative ‘Origin of the Universe’ and contract no. Ab128/2-1.

References

- [1] Collela R, Overhauser A W and Werner S A 1975 *Phys. Rev. Lett.* **34** 1472
- [2] Jenke T *et al* 2009 *Nucl. Instrum. Methods A* **611** 318
- [3] Abele H *et al* 2009 *Nucl. Phys. A* **827** 593
- [4] Reiter R, Schleder B and Seppi D 2009 *Quantenmechanisches Verhalten eines ultrakalten Neutrons im Gravitationsfeld* (Vienna: Projektarbeit, Atominstitut, Vienna University of Technology)
- [5] Nesvizhevsky V *et al* 2002 *Nature* **415** 297
- [6] Jenke T *et al* 2011 *Nature Phys.* **7** 468
- [7] Durstberger K *et al* 2011 *Phys. Rev. D* **84** 036004
- [8] Abele H *et al* 2010 *Phys. Rev. D* **81** 065019
- [9] Abele H 2008 *Prog. Part. Nucl. Phys.* **60** 1
- [10] Hasegawa Y and Rauch H 2011 *New J. Phys.* **13** 115010
- [11] Rauch H and Werner S A 2000 *Neutron Interferometry, Lessons in Experimental Quantum Mechanics* (Oxford: Oxford Science Publications, Clarendon)
- [12] Byrne J 1993 *Neutrons, Nuclei and Matter: An Exploration of the Physics of Slow Neutrons* (Bristol and Philadelphia: Institute of Physics Publishing)
- [13] Aminoff C G *et al* 1993 *Phys. Rev. Lett.* **71** 3083
- [14] Kasevich M A, Weiss D S and Chu S 1990 *Opt. Lett.* **15** 607
- [15] Roach T M *et al* 1995 *Phys. Rev. Lett.* **75** 629
- [16] McReynolds A W 1951 *Phys. Rev. Lett.* **83** 172
- [17] Dabbs J W T *et al* 1965 *Phys. Rev.* **139** B756
- [18] Koester L 1976 *Phys. Rev. D* **14** 907
- [19] Sears V F 1982 *Phys. Rep.* **82** 1
- [20] Rauch H, Treimer W and Bonse H 1974 *Phys. Lett. A* **47** 369
- [21] Werner S A *et al* 1988 *Physica B* **151** 22
- [22] Littrell K C *et al* 1997 *Phys. Rev. A* **56** 1767
- [23] Werner S A and Klein A G 1986 *Methods Exp. Phys. A* **23** 259
- [24] Atwood D K *et al* 1984 *Phys. Lett.* **52** 1673
- [25] Staudenmann J L *et al* 1980 *Phys. Rev. A* **21** 1419
- [26] van der Zouw G *et al* 2000 *Nucl. Instrum. Methods A* **440** 568
- [27] Mannheim P D 1998 *Phys. Rev. A* **57** 1260
- [28] Cohen J M and Mashhoon B 1993 *Phys. Lett. A* **181** 353
- [29] Cronin A D, Schmiedmayer J and Pritchard D E 2009 *Rev. Mod. Phys.* **81** 1051
- [30] Peters A, Chung K Y and Chu S 1999 *Nature* **400** 849
- [31] Müller H, Peters A and Chu S 2010 *Nature* **463** 927
- [32] Müller H, Peters A and Chu S 2010 *Nature* **467** E2
- [33] Wolf P *et al* 2010 *Nature* **467** E1
- [34] Wolf P *et al* 2011 *Class. Quantum Gravity* **28** 145017
- [35] Gibbs R L 1975 *Am. J. Phys.* **43** 25
- [36] Rosu H C 2001 arXiv:quant-ph/0104003
- [37] Gea-Banacloche J 1999 *Am. J. Phys.* **67** 776
- [38] Wallis H *et al* 1992 *Appl. Phys. B* **54** 407
- [39] Luschikov V I and Frank A I 1978 *JETP Lett.* **28** 559
- [40] Westphal A 2001 *Diploma Thesis* University of Heidelberg (arXiv:gr-qc/0208062)

- [41] Voronin Yu A 2006 *Phys. Rev. D* **73** 044029
- [42] Westphal A *et al* 2007 *Eur. Phys. J. C* **51** 367 (arXiv:hep-ph/0602093)
- [43] Ruess T 2000 *Diploma Thesis* University of Heidelberg
- [44] Nesvizhevsky V *et al* 2003 *Phys. Rev. D* **67** 102002
- [45] Nesvizhevsky V V *et al* 2005 *Euro. Phys. J. C* **40** 479
- [46] Rabi I *et al* 1939 *Phys. Rev.* **55** 526
- [47] Baker C A *et al* 2006 *Phys. Rev. Lett.* **97** 131801
- [48] Kreuz M *et al* 2009 arXiv:physics.ins-det/0902.0156
- [49] Hamilton W A *et al* 1987 *Phys. Rev. Lett.* **58** 2770
- [50] Felber J *et al* 1996 *Phys. Rev. A* **53** 319
- [51] Steane A *et al* 1995 *Phys. Rev. Lett.* **74** 4972
- [52] Bernet S *et al* 1996 *Phys. Rev. Lett.* **77** 5160
- [53] Arkani-Hamed N, Dimopoulos S and Dvali G 1999 *Phys. Rev. D* **59** 086004
- [54] Fischbach E and Talmadge C L 1999 *The Search for Non-Newtonian Gravity* (New York: Springer)
- [55] Callin P and Burgess C P 2006 *Nucl. Phys. B* **752** 60–79
- [56] Sundrum R 1999 *J. High Energy Phys.* JHEP07(1999)001
- [57] Kajari E *et al* 2010 *Appl. Phys. B* **100** 43–60
- [58] Jenke T 2011 *Dissertation Thesis* Vienna University of Technology