A photograph showing the interior of the ATLAS detector under construction. The image is a perspective view looking down a long, narrow tunnel. The tunnel is lined with complex machinery, including large cylindrical components with orange and white stripes. The structure is supported by a dense network of blue and green metal beams and scaffolding. The lighting is bright, highlighting the intricate details of the detector's construction.

# *Review di ricerche di SUSY in ATLAS*

**U. De Sanctis  
Università di Milano & INFN**



# Searches for Supersymmetry with ATLAS

- Search strategies for mSUGRA models
- Commissioning of the detector
- Measurement and control of backgrounds
- Other SUSY models and search strategies
- **After discovery**
  - Measurement of masses and other properties
  - From measurements to theory

# SUPERSYMMETRY REMINDER



Adds to each SM fermion (boson) a bosonic (fermionic) partner.

SM Particles	SUSY Particles	
quarks: $q$	$q$	squarks: $\tilde{q}$
leptons: $l$	$l$	sleptons: $\tilde{l}$
gluons: $g$	$g$	gluino: $\tilde{g}$
charged weak boson: $W^\pm$	$W^\pm$	Wino: $\tilde{W}^\pm$
Higgs: $H^0$	$H^\pm$ $h^0, A^0, H^0$	charged higgsino: $\tilde{H}^\pm$ neutral higgsino: $\tilde{h}^0, \tilde{A}^0$
neutral weak boson: $Z^0$	$Z^0$	Zino: $\tilde{Z}^0$
photon: $\gamma$	$\gamma$	photino: $\tilde{\gamma}$

$\left. \begin{array}{l} \tilde{W}^\pm \\ \tilde{H}^\pm \end{array} \right\} \tilde{\chi}_{1,2}^{\pm} \text{ chargino}$   
 $\left. \begin{array}{l} \tilde{h}^0, \tilde{A}^0 \\ \tilde{Z}^0 \\ \tilde{\gamma} \end{array} \right\} \tilde{\chi}_{1,2,3,4}^0 \text{ neutralino}$

- R-parity  $R = (-1)^{3(B-L)+2S}$  can be conserved (RPC) or violated (RPV)
- RPC implies:
  - SUSY particles produced in pairs
  - stable and neutral lightest SUSY particle (LSP)
  - no proton decay
- LSP is a good candidate for cold Dark Matter

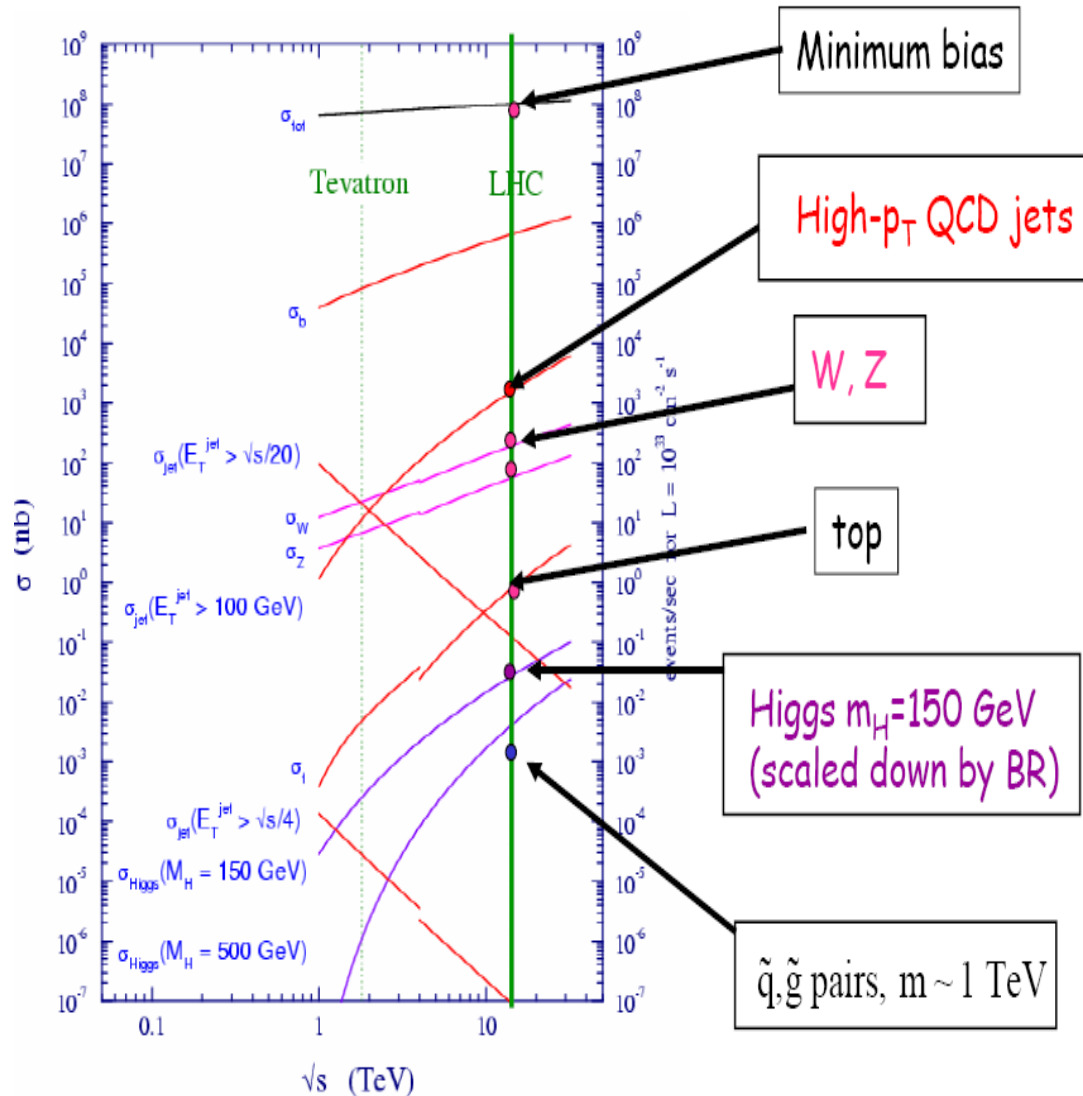
MSSM Lagrangian depends on 105 parameters  $\rightarrow$   
**mSUGRA** requires only 5 parameters  
 - Also other SUSY models exist: **GMSB**, **AMSB**, ...

Par.	Description
$m_0$	Common scalar mass
$m_{1/2}$	Common gaugino mass
$A_0$	Common trilinear term
$\tan\beta$	Ratio of Higgs vev
$\text{sign}(\mu)$	$\mu$ from Higgs sector

# A needle in an hay stack



Cross section (nb)



Only one event (i.e. pp collision) in **one billion** may contain an Higgs boson or a squark....

**Need high luminosity**

**Need an efficient online selection (trigger) to select interesting events:**  
cannot register everything electronically  
for further processing

# What do we do when we get the data?



Before we can claim discovery of “New Physics” we have to do some homework...

- Understand and calibrate detector and trigger in situ using well-known physics samples:  $Z/W \rightarrow$  leptons, semileptonic  $tt$
- Understand basic SM physics at 14 TeV: first measurements and publications
  - jets and  $W, Z$  cross-section top mass and cross-section
  - Event features: Min. bias, jet distributions, PDF constraints
- Understand tails of SM processes as backgrounds ( $tt, W/Z +$  jets), go for discovery:  $Z',$  SUSY, Higgs

But let's have a look at our main SUSY discovery strategy, to understand what we need to understand to get there...

# *mSUGRA benchmark points*



SUSY **benchmark points** chosen in the  $(m_0, m_{1/2})$  plane for different  $\tan\beta$  values:

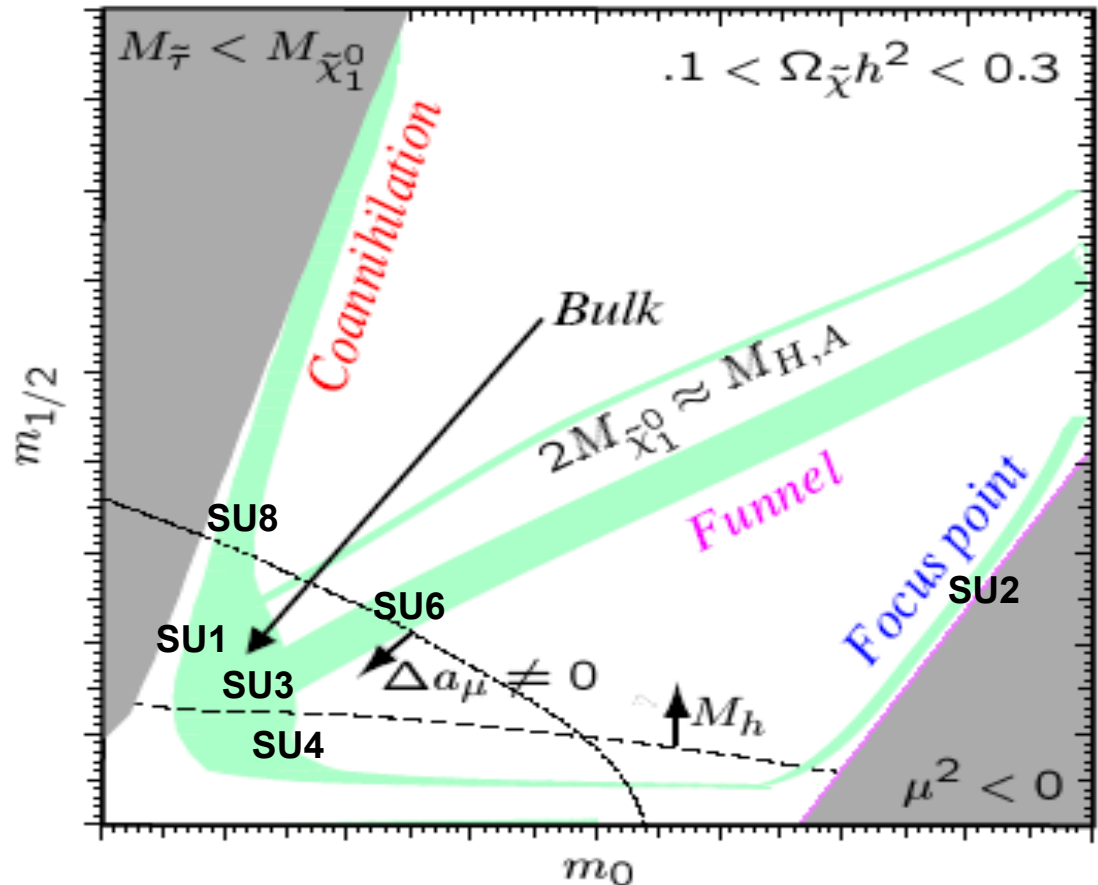
- ✓ Systematically exploring phenomenological signatures
- ✓ Scanning the parameter phase space constrained by latest experimental data

*Coannihilation:* Light  $\tilde{\tau}_1$  in equilibrium with  $\tilde{\chi}_1^0$ , so annihilate via  $\tilde{\chi}_1^0 \tilde{\tau}_1 \rightarrow \gamma \tau$ .

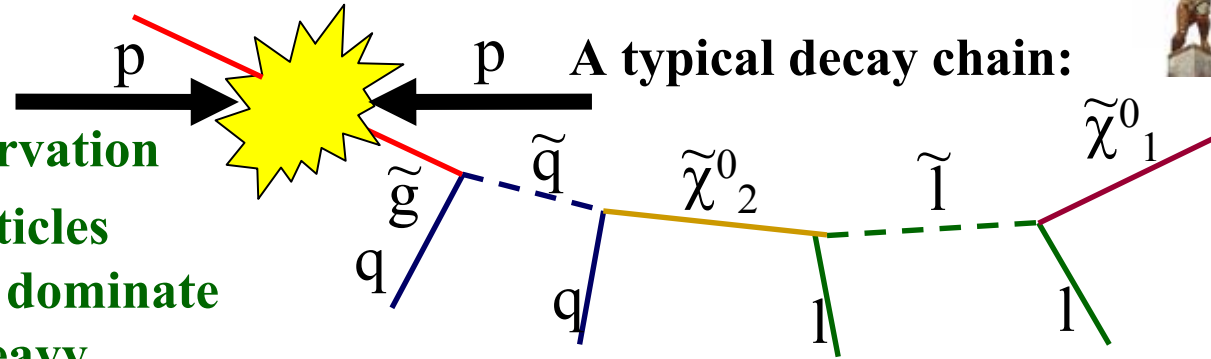
*Bulk:* bino  $\tilde{\chi}_1^0$ ; light  $\tilde{\ell}_R$  enhances annihilation.

*Funnel:*  $H, A$  poles enhance annihilation for  $\tan\beta \gg 1$ .

*Focus point:* Small  $\mu^2$ , so Higgsino  $\tilde{\chi}_1^0$  annihilate. Heavy s-fermions, so small FCNC.

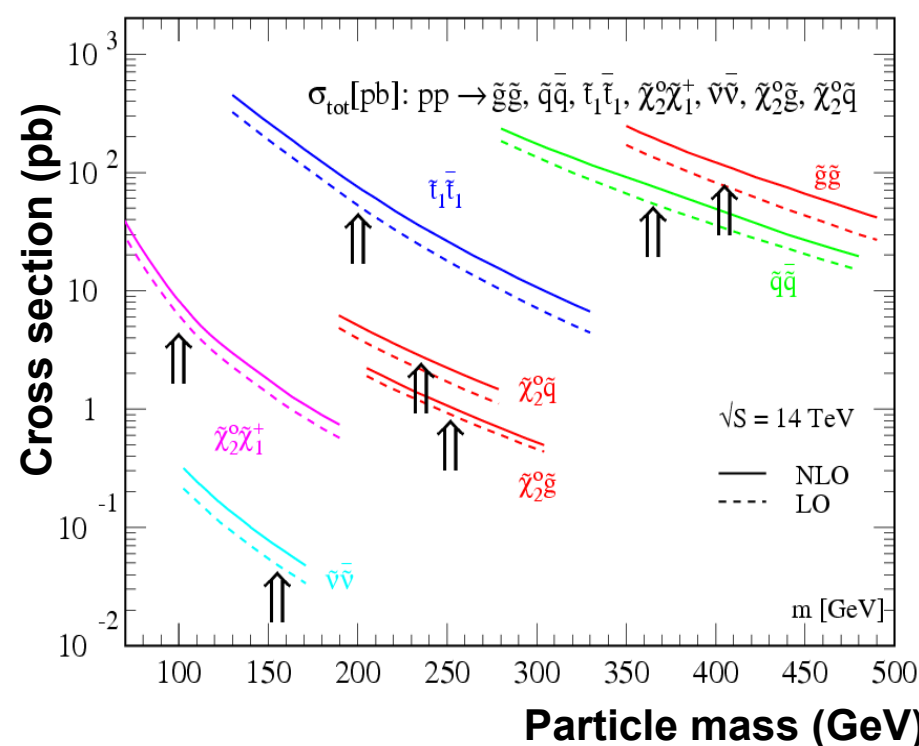


# SUSY signatures at an hadronic collider



- Assuming R-parity conservation
- Strongly interacting sparticles (squarks, gluinos) should dominate production unless very heavy.
- Cascade decays to the stable, weakly interacting lightest neutralino follows.
- Event topology:
  - high  $p_T$  jets (from squark/gluino decay)
  - Large  $E_T^{\text{miss}}$  signature (from LSP)
  - High  $p_T$  leptons, b-jets,  $\tau$ -jets (depending on model parameters).

Several other possibilities exist, some are mentioned later in this talk, but our effort has to be as more “model independent” as possible.



# Event topologies and baseline selection



Early searches try to cover a broad range of experimental signatures, but they are classified based on the event topology:

Large  $E_T^{\text{miss}}$  +

Jet multiplicity	Additional signature	SUSY scenario	Backgrounds
$\geq 4$	No lepton	<i>mSUGRA, AMSB, split SUSY, heavy squark</i>	QCD, <i>ttbar</i> , W/Z
	One lepton ( <i>e, μ</i> )	<i>mSUGRA, AMSB, split SUSY, heavy squark</i>	<i>ttbar</i> , W
	di-lepton	<i>mSUGRA, AMSB, GMSB</i>	<i>ttbar</i>
	di-tau	<i>GMSB, large tan β</i>	<i>ttbar</i> , W
	$\gamma\gamma$	<i>GMSB</i>	free
$\sim 2$		<i>light squark</i>	Z

## Baseline selection (to be optimized)

- Jet multiplicity  $\geq 4$ ,  $p_T^{\text{1st}} > 100\text{GeV}$ ,  $p_T^{\text{others}} > 50\text{GeV}$
- $E_T^{\text{miss}} > \max(100\text{GeV}, 0.2 \times M_{\text{eff}})$
- Transverse sphericity  $> 0.2$
- Additional cuts depending on signature: Transverse mass  $> 100\text{GeV}$ ,  $p_T^{\text{lepton}} > 20\text{GeV}$  ( for one-lepton mode), harder cuts on  $M_{\text{eff}} \dots$



# SUSY search strategies



Most promising search strategy:  
jets +  $E_T^{\text{miss}}$  + n-leptons

- Real missing energy from SM processes with hard neutrino (tt, W+jets, Z+jets, bb\*, cc\*)

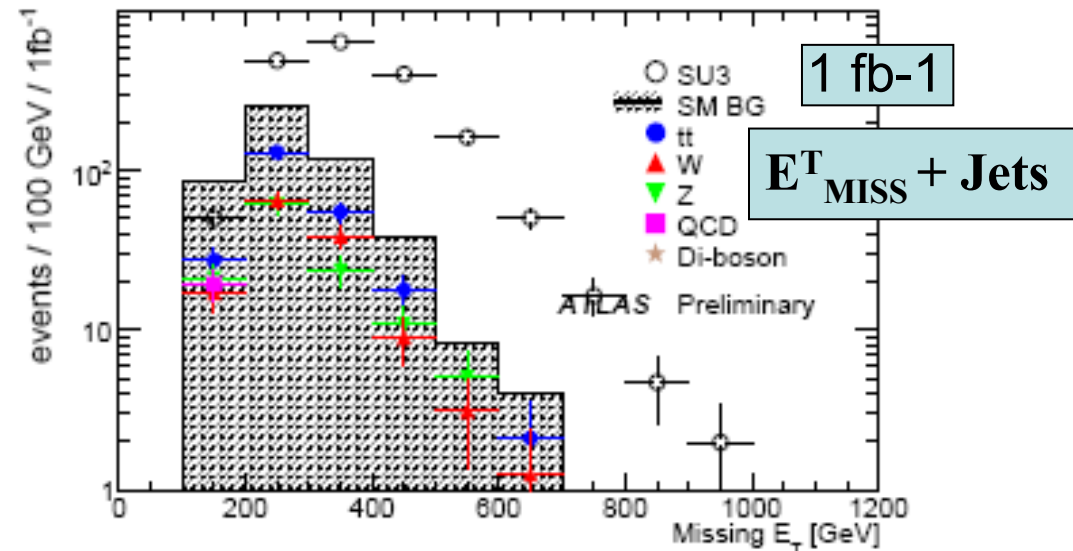
\*  $\nu$  from semileptonic B/D decay

- Fake missing energy from detector

Jet energy resolution (especially non-gaussian tails) critical

A good understanding of both SM physics and detector (missing energy especially) critical to claim excess over SM predictions

$E_T^{\text{MISS}}$  distribution after baseline selection  
(additional cut on  $M_{\text{eff}} > 800$  GeV)



SUSY → HERWIG + Isajet (for mass spectrum)

ttbar → MC@NLO

W,Z + jets → ALPGEN (1° jet > 80 GeV, 4° jet > 40 GeV, MET > 80 GeV)

QCD → PYTHIA (>=2 jets, 1° jet > 80 GeV, 2° jet > 40 GeV, MET > 100 GeV)

WW,ZZ,WZ → HERWIG

MCWS

U. De Sanctis

# Detector calibration and alignment



The jet energy scale affects directly SUSY discovery plots through the cut on the presence of hard jets.

Also,  $E_T^{\text{miss}}$  depends on the correct reconstruction of the energies of jets, photons, electrons, and muons!

- We will start from the knowledge obtained from test-beam data, electronics calibrations, survey measurements during installation of the tracking detectors, and cosmics data.
- We will then use well-known SM processes (standard candles) to improve

## Examples: leptonic decays of Z, W mass in semileptonic top events

	Expected performance day-1	Physics samples to improve (examples)
ECAL uniformity e/ $\gamma$ E-scale	1-2% (~0.5% locally) ~ 2 %	Isolated electrons, $Z \rightarrow ee$ $Z \rightarrow ee$
HCAL uniformity	~ 3 %	Single pions, QCD jets
Jet E-scale	< 10%	$\gamma/Z + 1j$ , $W \rightarrow jj$ in $t\bar{t}$ events
Tracking alignment	10-200 $\mu\text{m}$ in $R\phi$ Pixels/SCT ?	Generic tracks, isolated $\mu$ , $Z \rightarrow \mu\mu$

Process	$\sigma \times BR$	<b>Eff.</b>	Events selected for $100 \text{ pb}^{-1}$
$W \rightarrow \ell\nu$	20 nb	~ 20%	~ 400000
$Z \rightarrow \mu\mu$	2 nb	~ 20%	~ 40000
$t\bar{t}$ (semileptonic)	370 pb	~ 1.5%	< 1000

Available statistics, with conservative estimates of reconstruction efficiencies

# $E_T^{\text{miss}}$ Commissioning: Event cleaning



Raw  $E_T^{\text{miss}}$  in early data is expected to have large tails

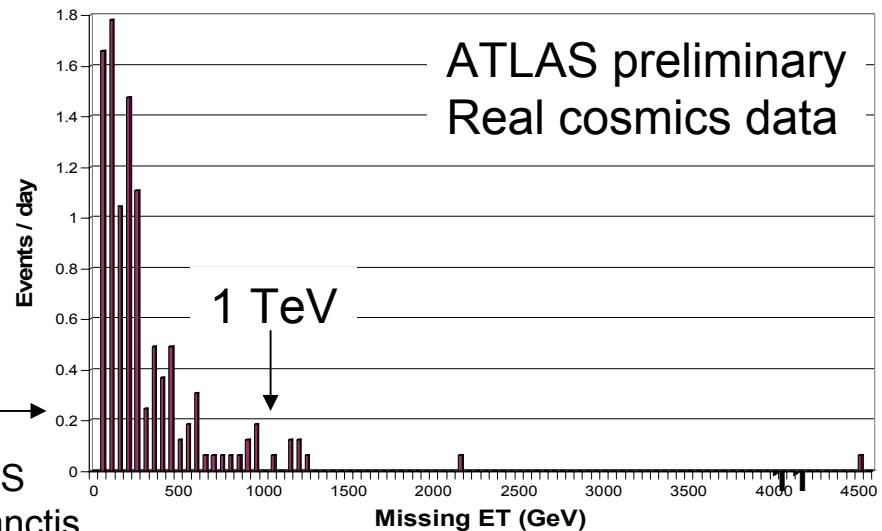
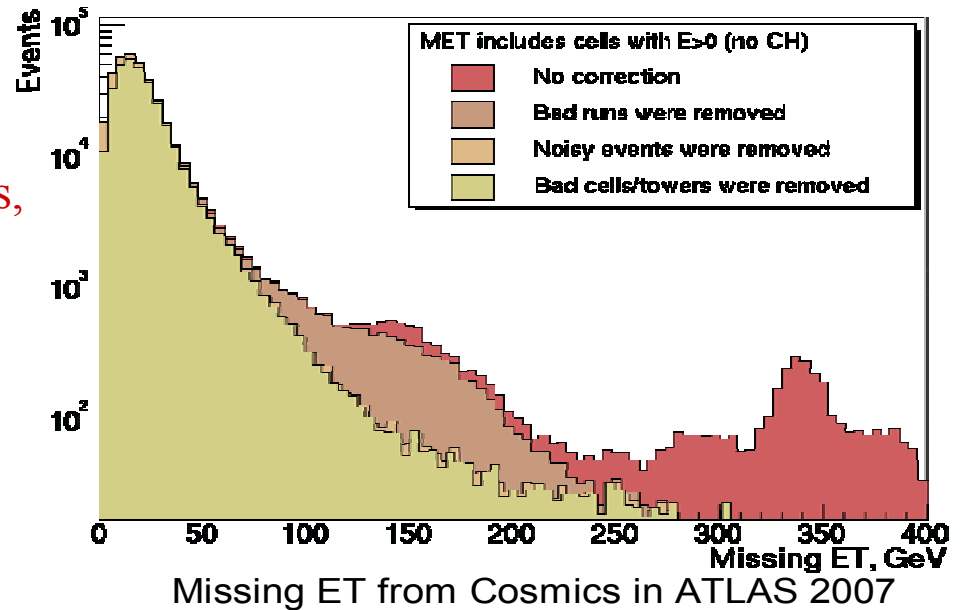
- Cosmic events
- Beam halo muons, beam gas interactions, cavern background (neutrons)
- Noisy and dead calorimeter cells

All machine and detector garbage collected by  $E_T^{\text{miss}}$  trigger!

We are developing tools for event cleaning

- Online and offline monitoring
- Detect noisy/dead cells
- Reject beam halo and cosmic events
- $E_T^{\text{miss}}$  correlation with hardest jet, muons,
- Stability of  $E_T^{\text{miss}}$  trigger rate

## Effect of event cleaning on D0 $E_T^{\text{miss}}$



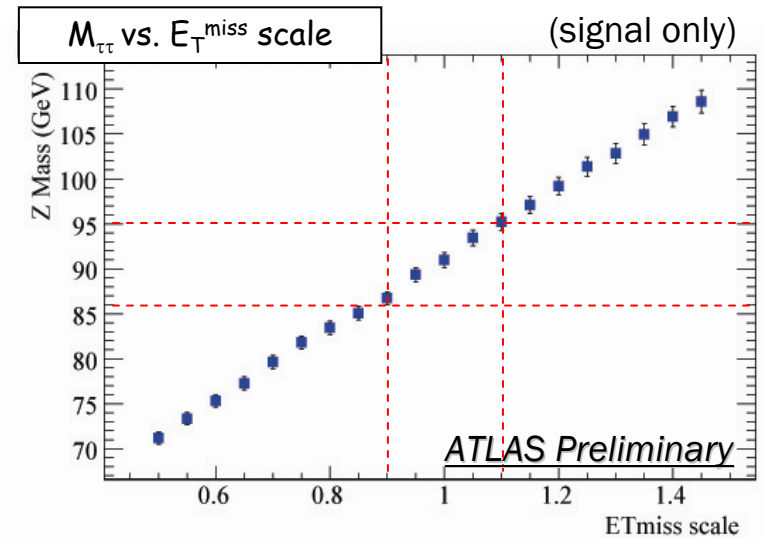
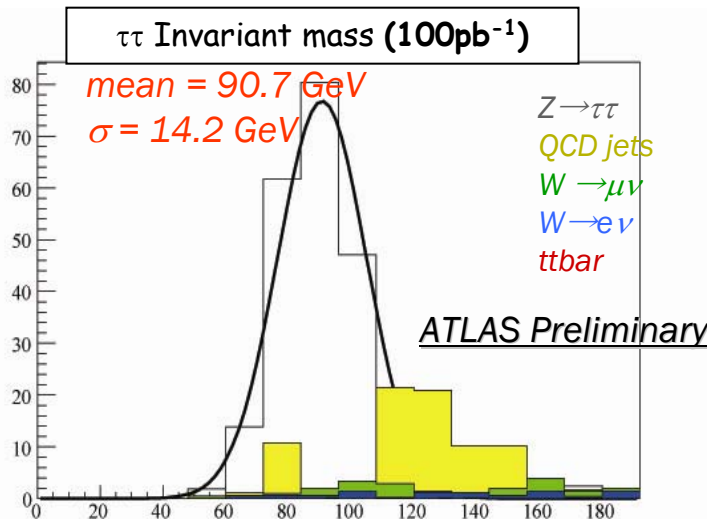
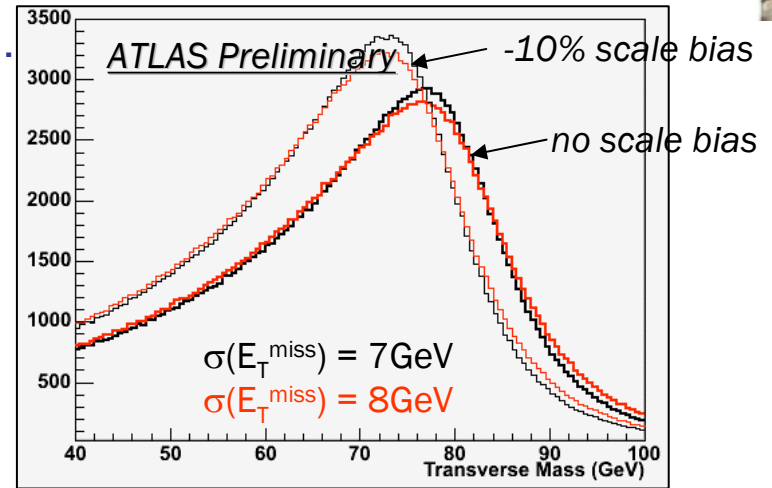
# $E_T^{MISS}$ Commissioning (2)



Just two examples, several other physics process can be used: minimum bias, Z(ll), ttbar, ...

**W(lv) sample:** Shape of transverse mass distribution depends on  $E_T^{miss}$  scale and resolution.

**Z( $\tau\tau$ ) sample:** Z mass can be reconstructed with collinear approximation (since the  $\tau$  are boosted,  $\nu$  are along visible  $\tau$  energy). Can be used to calibrate  $E_T^{miss}$  scale.



10% in  $E_T^{miss}$  scale  $\Leftrightarrow$  3% shift in Z mass

## *Towards SUSY searches...*



Once detector effects are understood, the next steps are:

- Fiducial cuts: reject  $E_T^{\text{miss}}$  pointing along leading jets, events with jets or electrons in calorimeter crack...
- Measure Z, W, ttbar cross sections and PDFs
- Understand residual tails in  $E_T^{\text{miss}}$  performance and distribution of real  $E_T^{\text{miss}}$  in SM events

Use data-driven estimates, do not rely on MC predictions

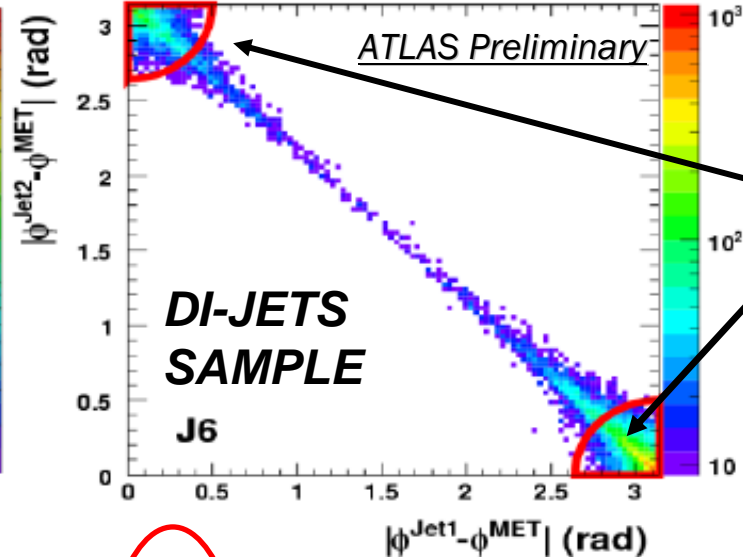
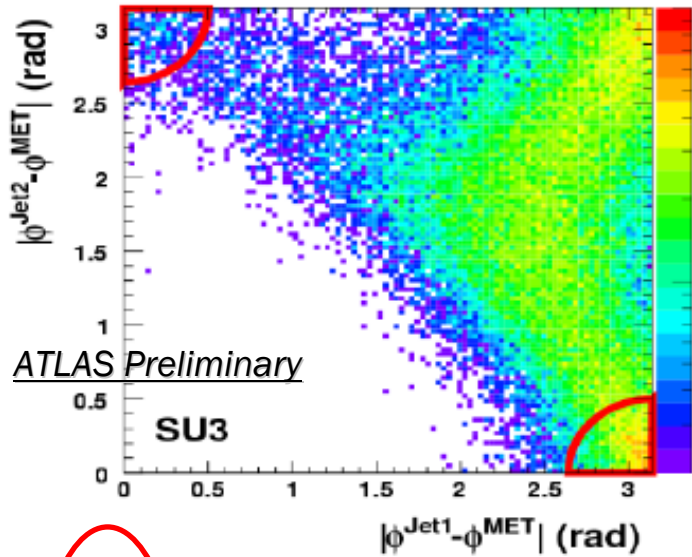
Some examples in next slides, several other techniques are being studied. Results should be available early next year

The aim is to estimate the background for each channel with at least two independent technique and compare the results to get confidence that we really understand the SM background

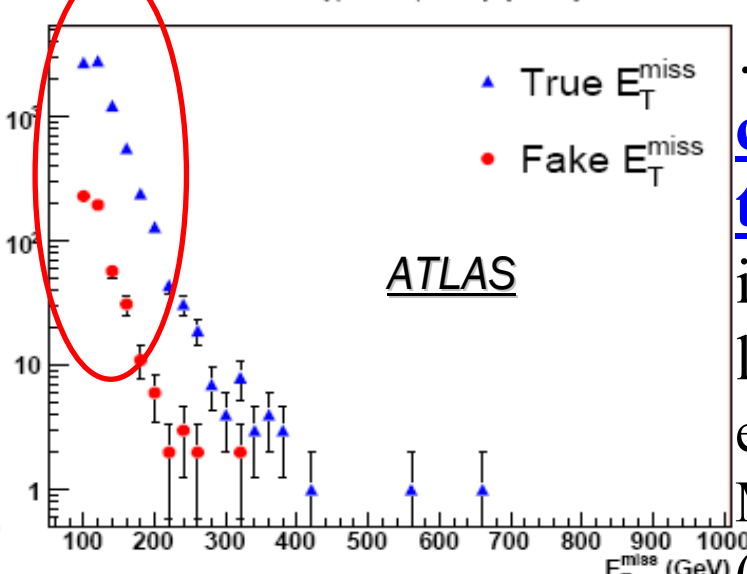
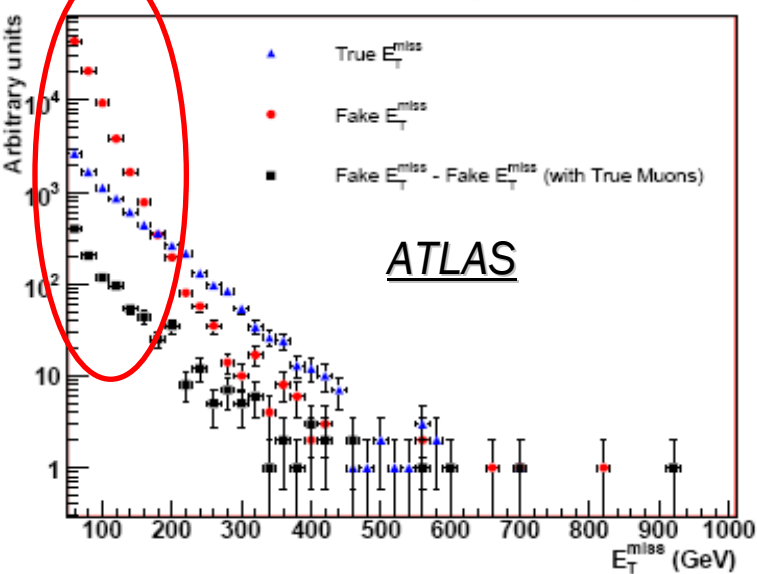
# $E_T^{MISS}$ & QCD (1)



Azimuthal difference of 2 leading jets w.r.t. MET vector



Avoiding these regions.....



..... QCD contribution to fake MET

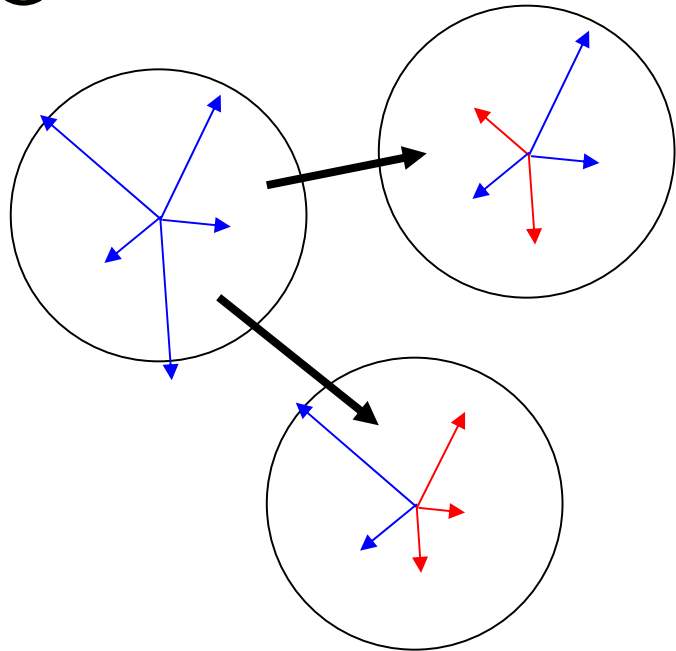
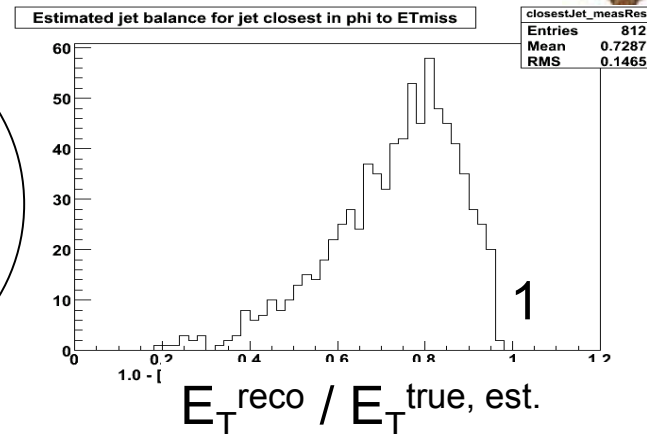
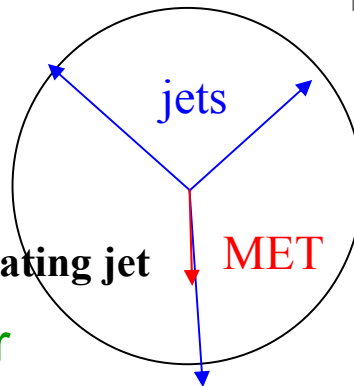
is strongly lowered, especially for MET < 200 GeV (red circles).

# $E_T^{\text{MISS}} \& \text{QCD (2)}$



① Measure smearing function in events with large  $E_T^{\text{miss}}$

② Select seed events and smear

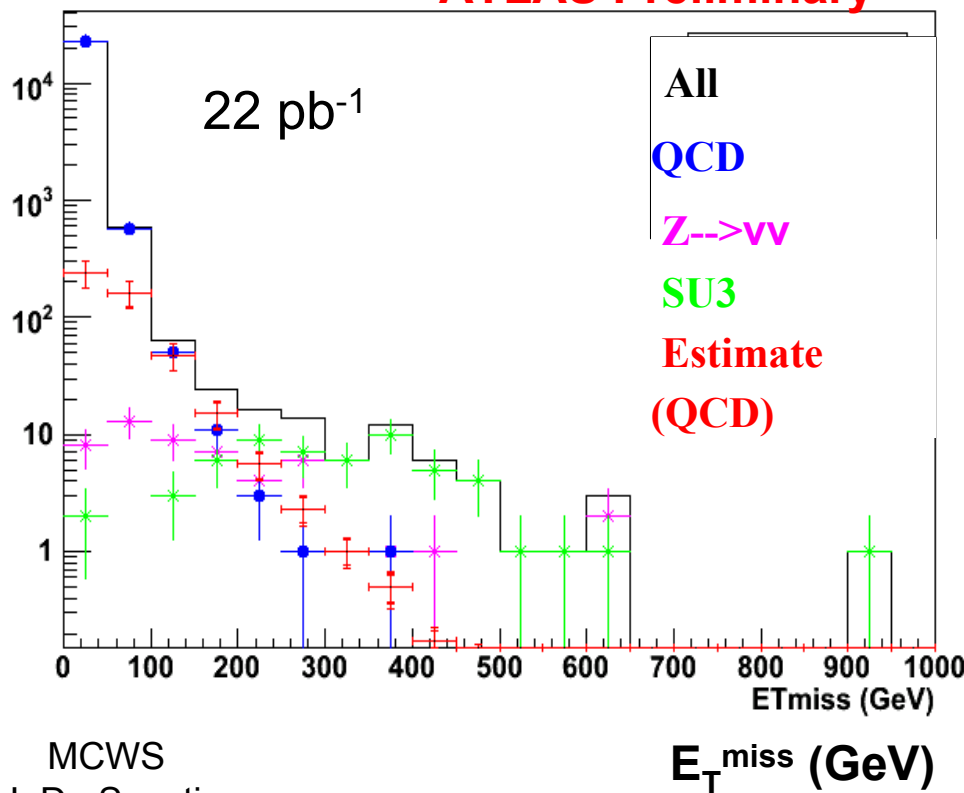


Seed events: low  $E_T^{\text{miss}} / \sum E_T$

③ Normalize estimate to data

ETmiss (inclusive)

ATLAS Preliminary

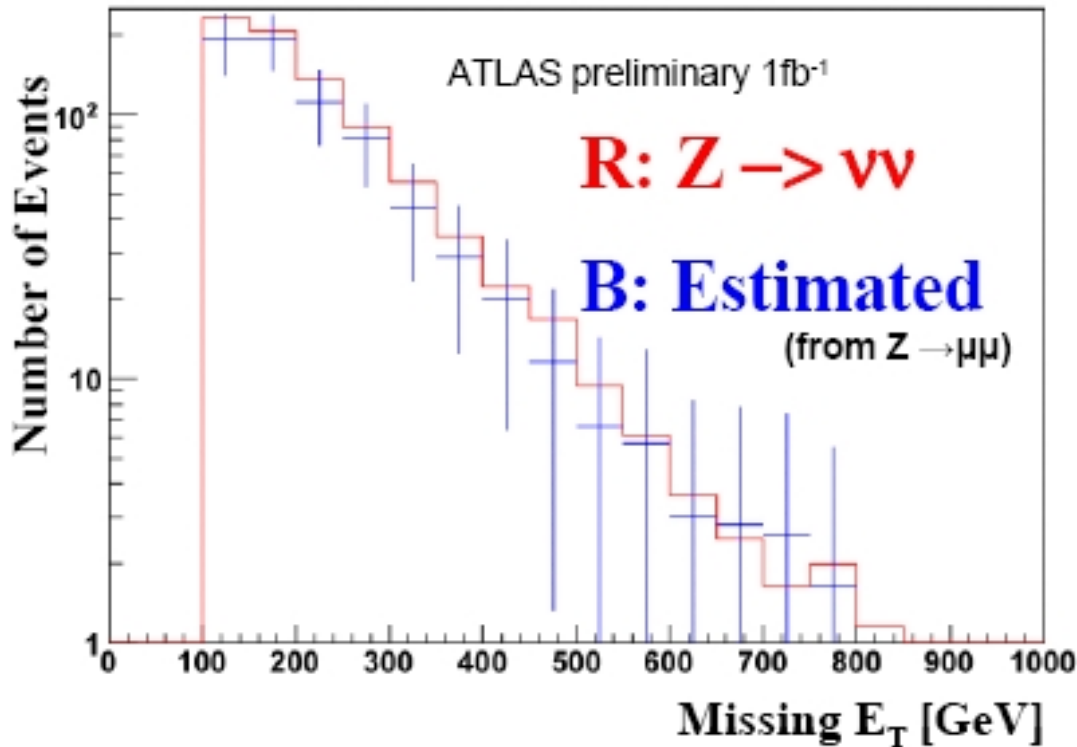


# *Z and W background*



Estimate  $Z \rightarrow \nu\nu$  from  $Z \rightarrow \ell^+\ell^-$

0-lepton mode: Z+jets



Similarly we can use the Z distribution to estimate the  $W \rightarrow \ell\nu$  background



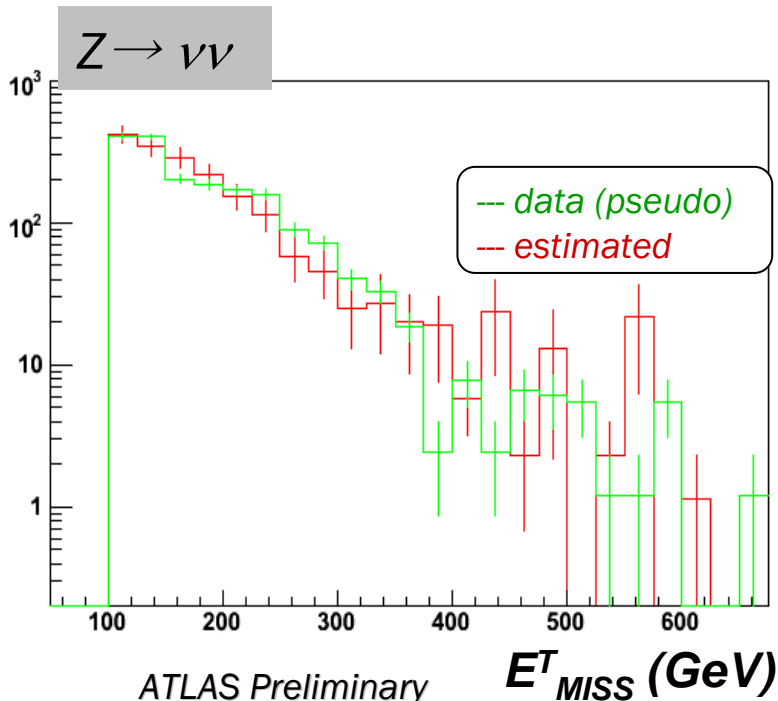
# Z and W background (0-lepton mode)



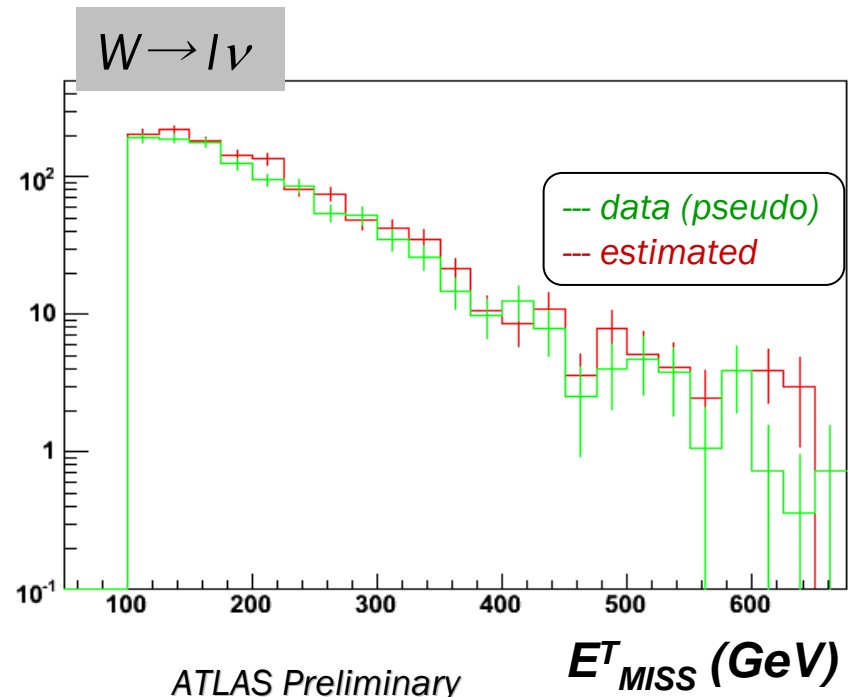
$Z \rightarrow \nu\nu$  and  $W \rightarrow l\nu$  can be estimated from  $Z \rightarrow \ell^+\ell^-$

Either **replace** the two leptons with neutrinos correcting for acceptance and efficiency

Or determine the **MC normalization** from  $Z(l\ell)$  and apply it to normalize the MC distribution of  $Z(\nu\nu)$  and  $W(l\nu)$  (almost same production mechanism)



Frascati, 19/02/2008



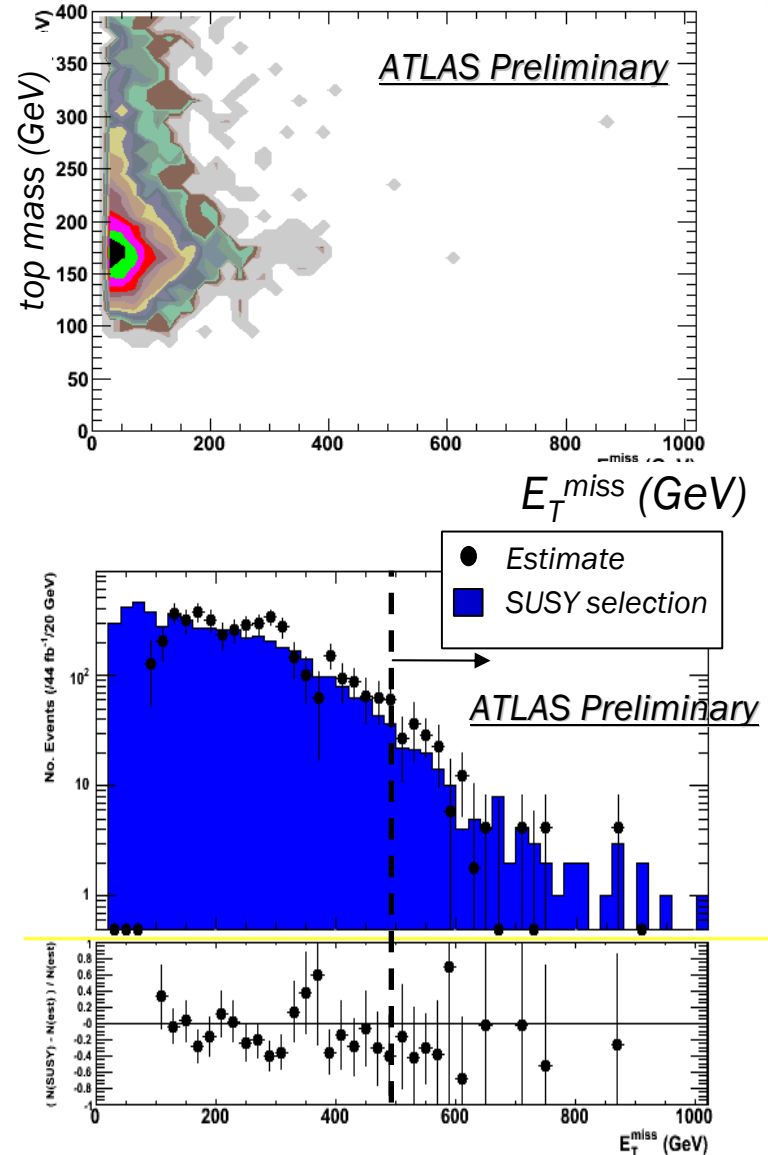
MCWS  
U. De Sanctis

# *ttbar background*

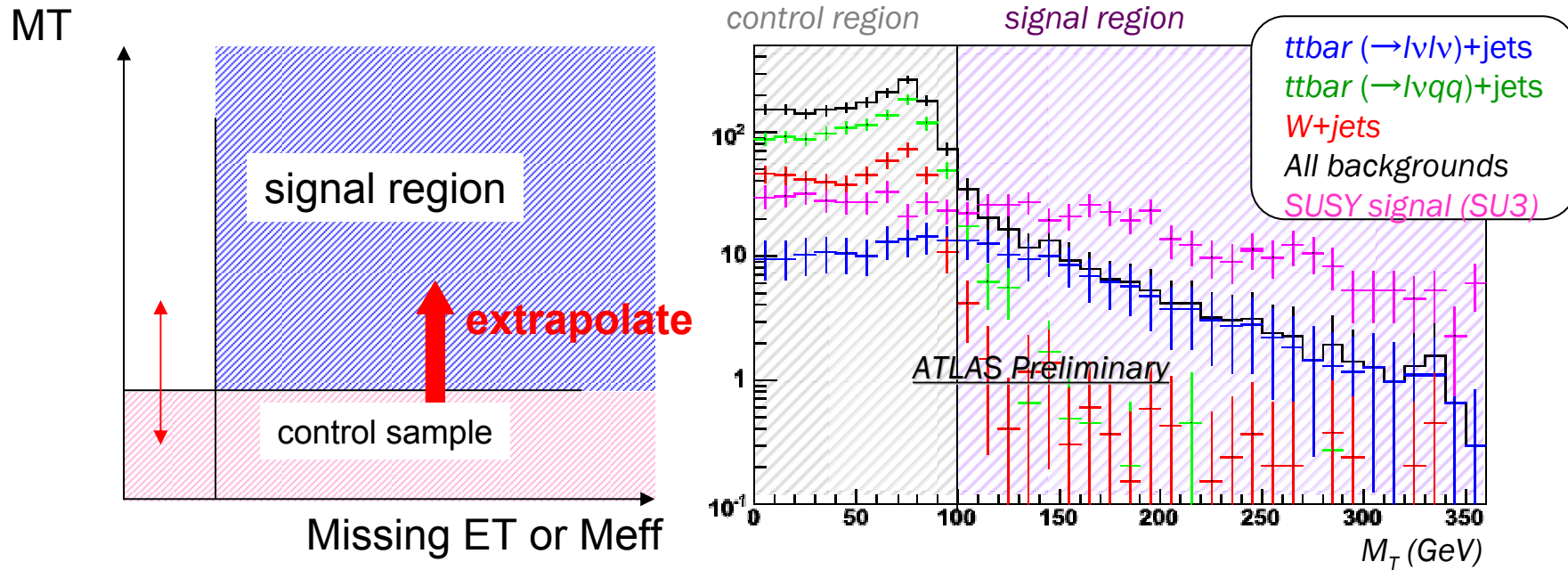


1. Top mass is largely uncorrelated with  $E_T^{\text{miss}}$ 
  - used as a calibration variable
2. Select semi-leptonic top candidates
  - mass window: 140-200 GeV
3. Contributions of combinatorial BG to top mass are estimated from the side-band events ( $200\text{GeV} < m_{\text{top}} < 260\text{GeV}$ )
4. Normalize the  $E_T^{\text{miss}}$  distribution in low  $E_T^{\text{miss}}$  region where SUSY signal contamination is small.
5. Extrapolate it to high  $E_T^{\text{MISS}}$  region and estimate the background with SUSY signal selection.

Several other techniques also under investigation



# Transverse mass method



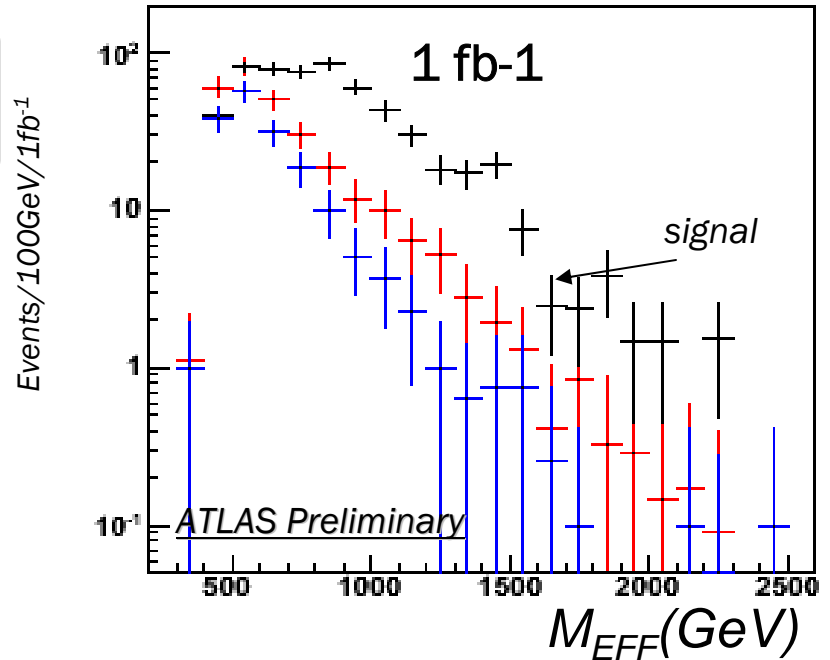
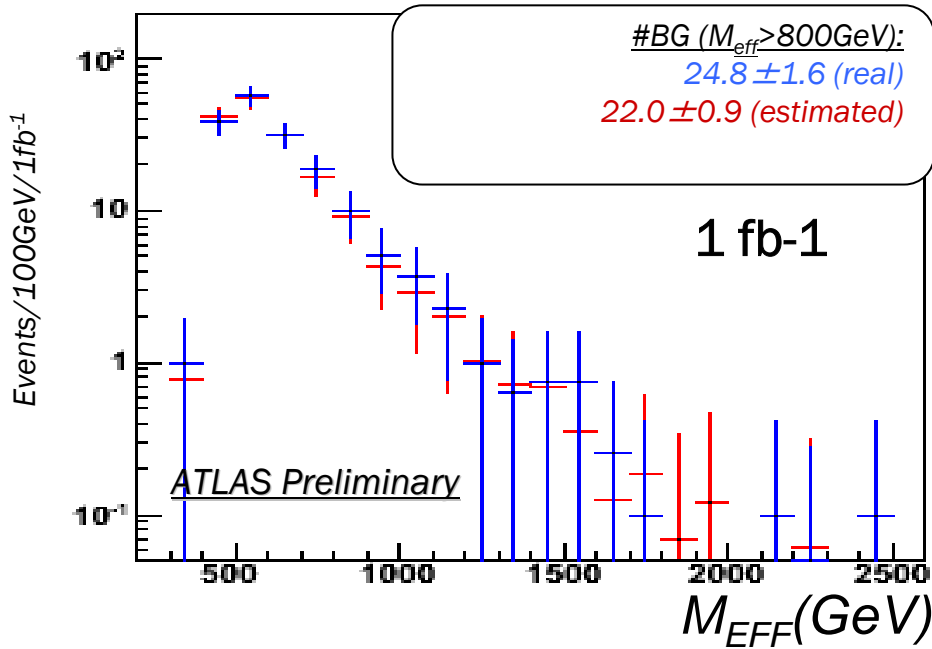
1. Define control sample with transverse mass  $< 100 \text{ GeV}$
2. Estimate the  $E_T^{\text{miss}}/M_{\text{eff}}$  shapes of background processes using control sample
3. Determine the normalization of backgrounds with low  $E_T^{\text{miss}}$  regions of control and signal samples.

**Can be used for both W and top backgrounds in 0-lepton, 1-lepton and 2 lepton channels (results shown here for 1-lepton)**



# Transverse mass method

Including SUSY signal (SU3)



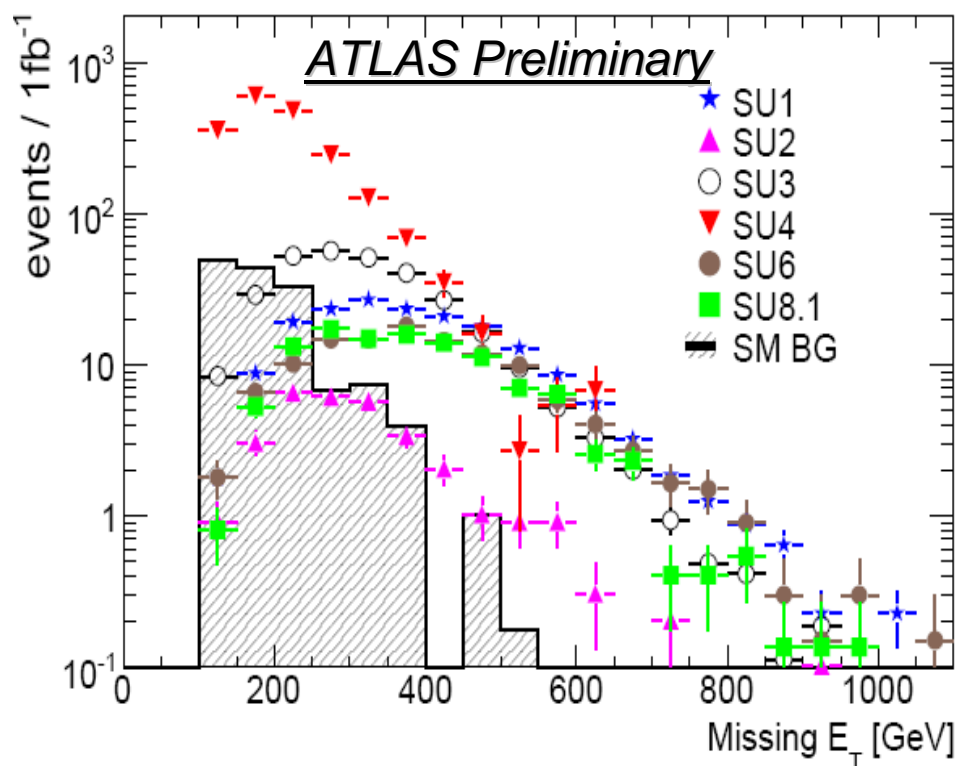
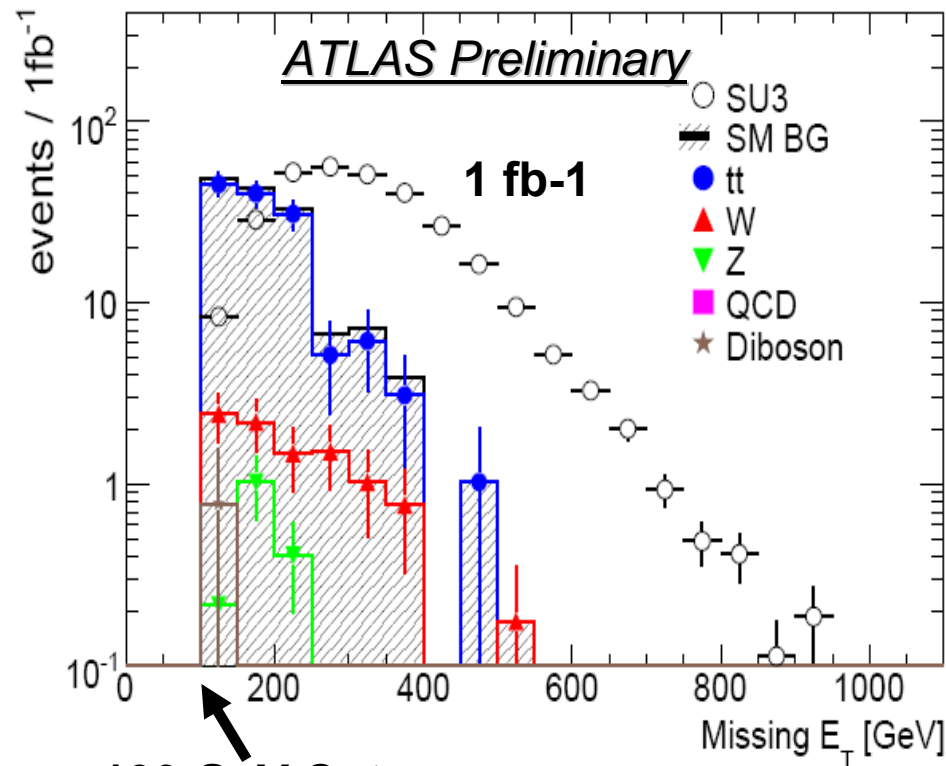
- **Satisfying performances** with the  $M_T$  discrimination technique.
- However, taking account of SUSY signal contamination in the control sample, this estimate appears to be over the mark (by a factor of 2.5 for SU3).  
It would not prevent discovery.

# Other strategy: 1-lepton channel



## Removing the lepton-veto: 1 lepton + Jets + $E_{\text{MISS}}^T$ channel

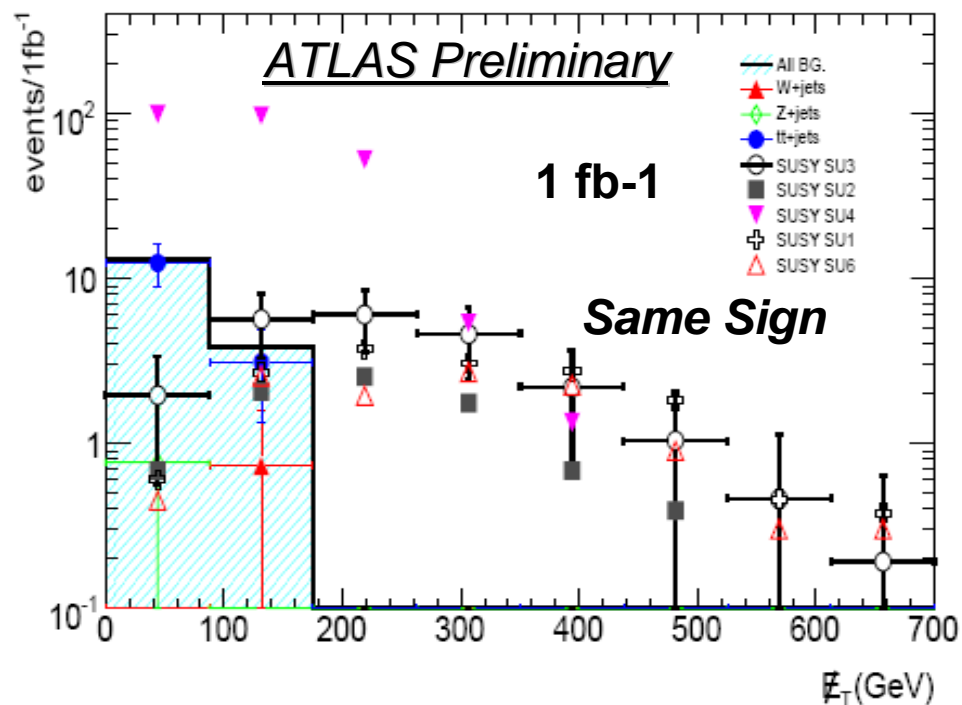
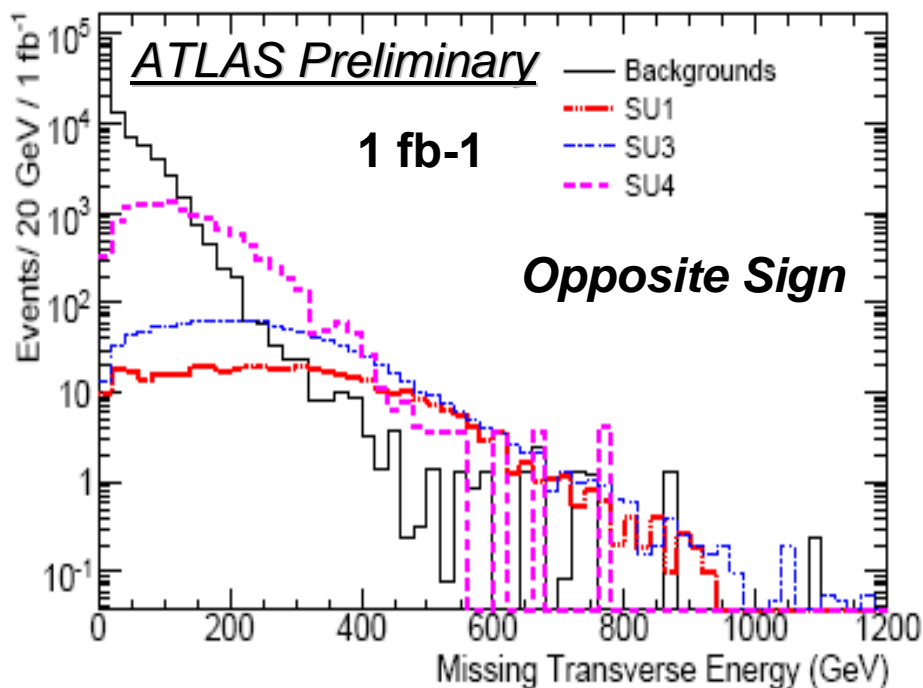
- Lepton can usually come from chargino/neutralino decays into LSP
- Heavily suppression of the QCD background (difficult to estimate from data and also with MC) requiring 1 isolated lepton with  $P_T > 20 \text{ GeV}/c$ .
- Dominant background are the same as the 0-lepton channel (except QCD) : top seems to be the dominant one, but  $W + \text{jets}$  is not negligible.



# Other strategy: 2-lepton channel



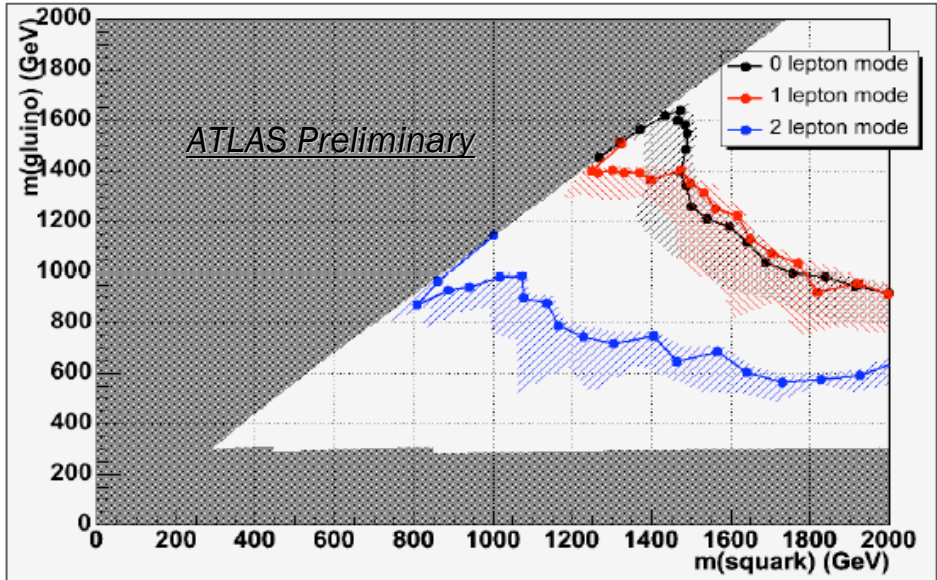
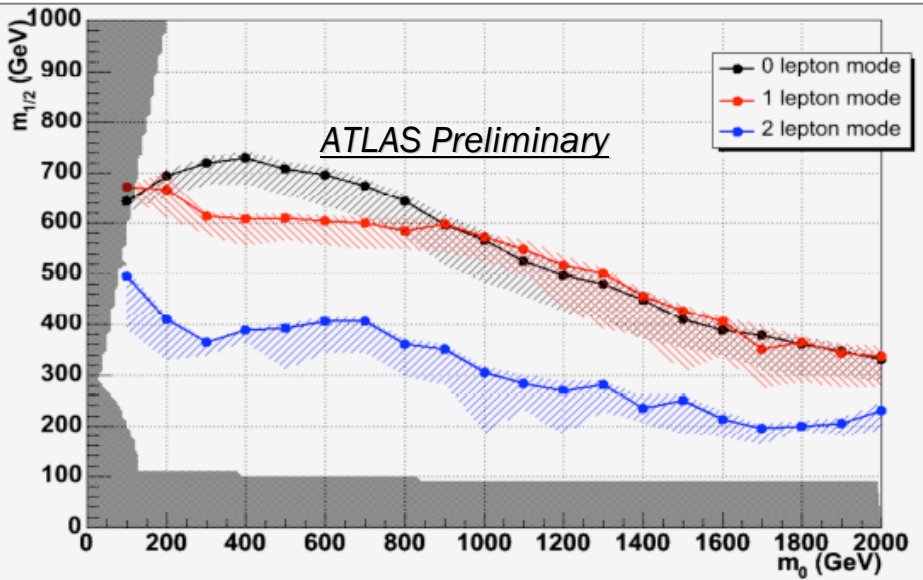
- Increasing the number of leptons
  - Reduces the signal because of (model dependent) leptonic BRs
  - Heavily suppresses the background
  - Statistical significance is smaller but S/B ratio larger. Top is dominant background
  - The Same Sign channel has the best S/B ratio – but limited by signal rate





# Back to SUSY discovery

- When the detector performance and the 14 TeV SM physics will be understood, we will be able to use the full power of our experiments for SUSY searches.
- Hopefully, we will still have an excess....



*5 $\sigma$ -discovery potential on  $m_{1/2}$ - $m_0$  ( $m_{gluino}$ - $m_{squark}$ ) space is shown for  $1 \text{ fb}^{-1}$   
Require  $S > 10$  and  $S/\sqrt{B} > 5$   
Factor of 2 generator-level uncertainty included (hatched)*

# What next ?

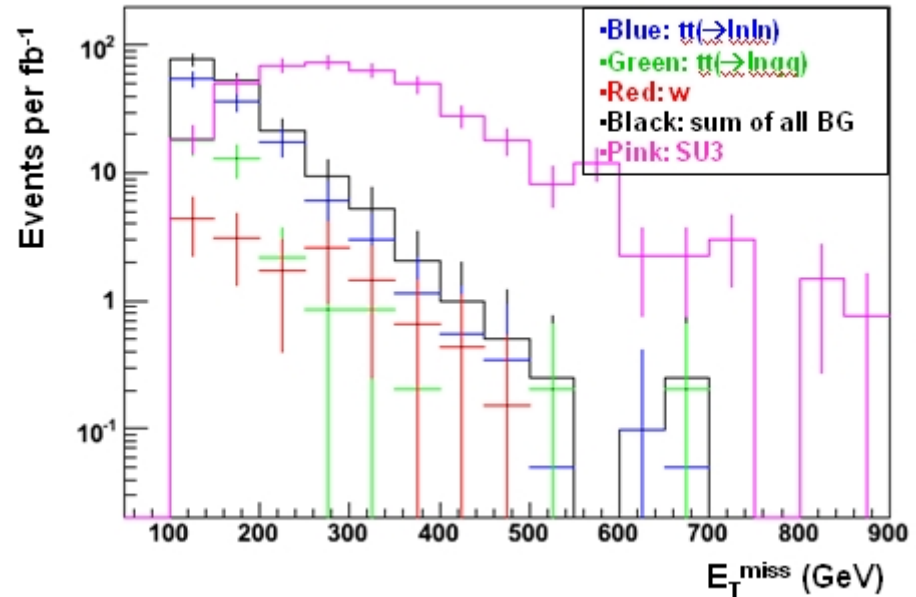
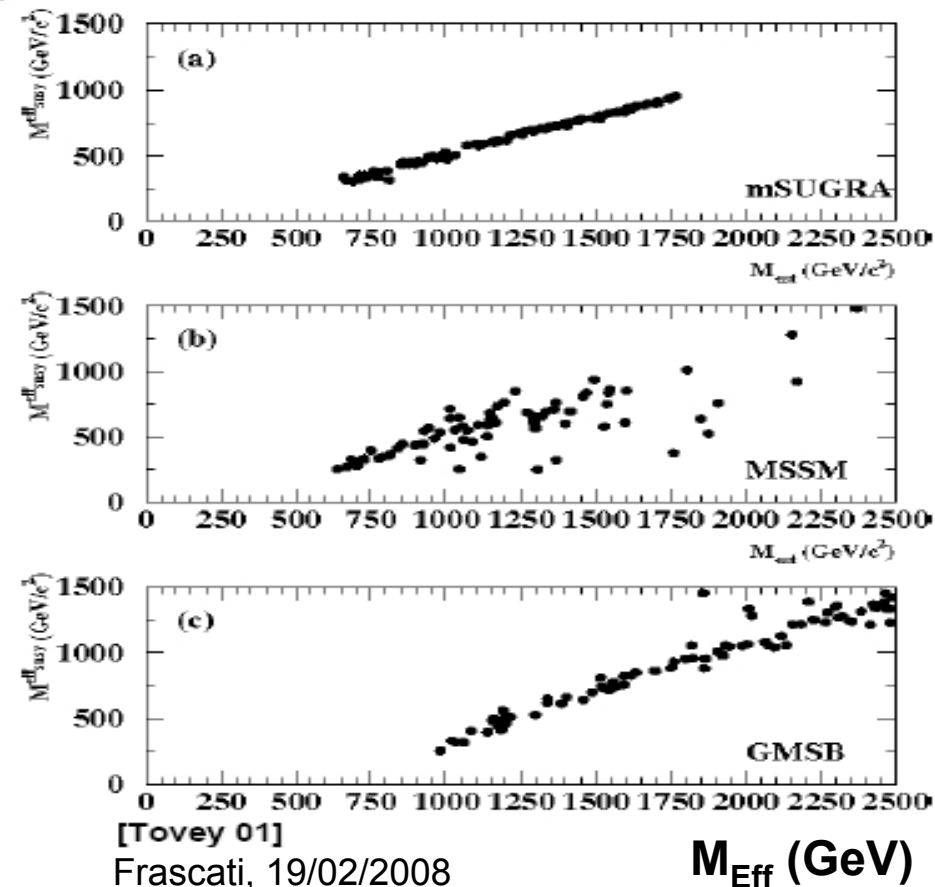


“Observation of an excess of events in multijet+MET events in pp collisions at 14 TeV with the ATLAS detector”

Large (>100GeV) Missing ET events:  
Smoking gun of Supersymmetry

## Is it SUSY?

If yes, what are the model parameters?



Measurement of the “effective mass” peak correlates with the SUSY mass scale (average squark, gluino mass)

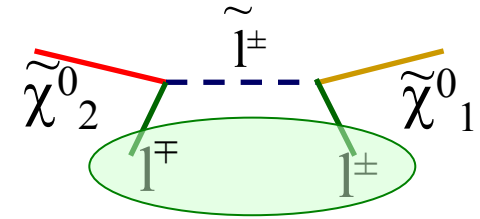
$M_{\text{eff}} = \text{MET} + \text{PT}_{1,2,3,4}$   
15% (40%) precision on  $M(\text{SUSY})$   
with 10fb<sup>-1</sup> for mSUGRA (MSSM)



# Di-Lepton Edge mass measurement (1)



- In case of a discovery of SUSY, **particle properties** can be measured to verify that they are indeed **SUSY partners**
- Edge(s) of **di-lepton invariant mass** correlated with slepton and neutralino masses
- Impossible to reconstruct peaks because  $\tilde{\chi}_1^0$  (LSP) escapes detection, more complicated relations between masses of particles involved.



$$\tilde{\chi}_2^0 \rightarrow \tilde{l} l \rightarrow \tilde{\chi}_1^0 l^+ l^-$$

$$M_{ll}^{\max} = M(\tilde{\chi}_2^0) \sqrt{1 - \frac{M^2(\tilde{l}_R)}{M^2(\tilde{\chi}_2^0)}} \sqrt{1 - \frac{M^2(\tilde{\chi}_1^0)}{M^2(\tilde{l}_R)}}$$

- ✓ Uncorrelated (SUSY+SM) **background** (two leptons from independent chains) **removed** by **flavour subtraction**:

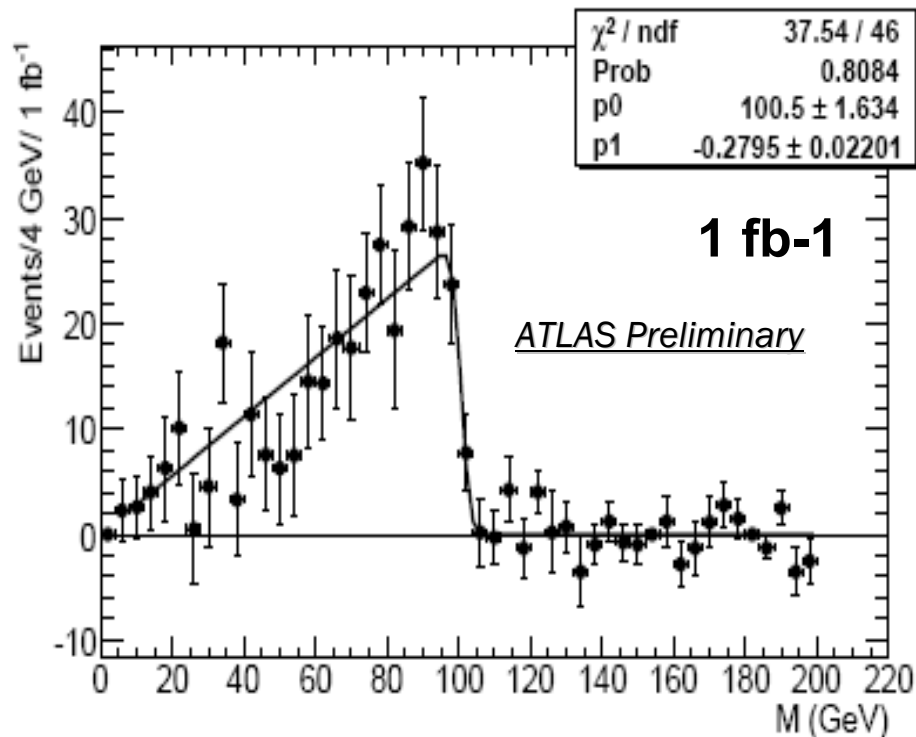
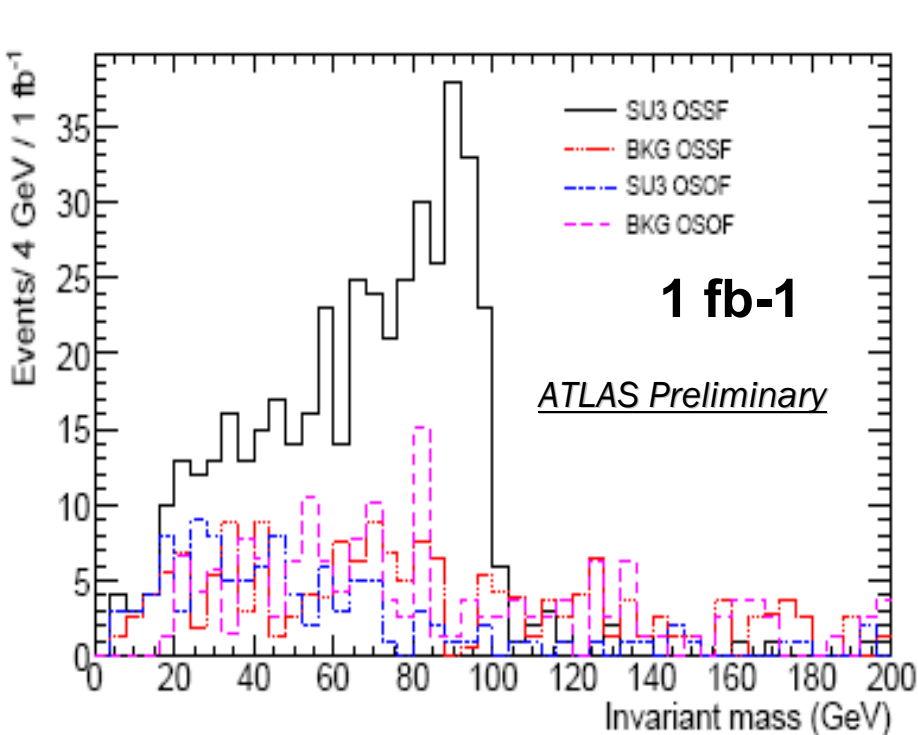
$$e^+e^- + \beta^2 \mu^+\mu^- - \beta (e^+\mu^- - e^-\mu^+), \quad \beta = \epsilon_e / \epsilon_\mu$$

- ✓ Leptons can also be combined with jets of the full decay chain to look for other **kinematical edges** ( $M_{llj}$  or  $M_{lj}$ )

# Di-Lepton Edge mass measurement (2)



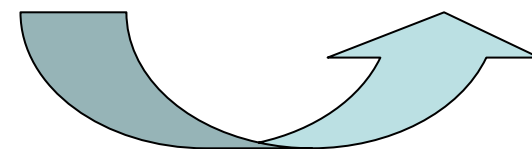
## Flavour subtraction at work....



**SU3, 1 fb<sup>-1</sup>**

Edge: (100.5±1.6) GeV

Truth: 100.2 GeV

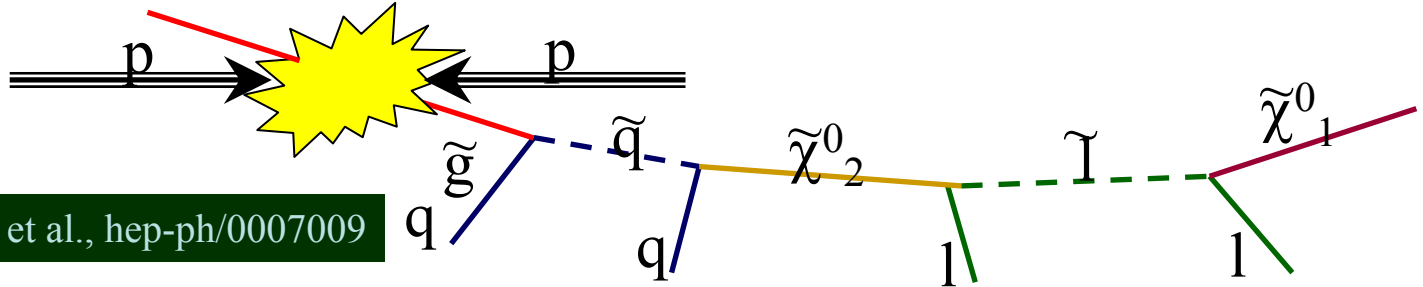


**Flavour Subtraction**

**Fitting function:**

Triangle smeared with a Gaussian with  $\sigma = 2$  GeV (to take into account experimental resolution)

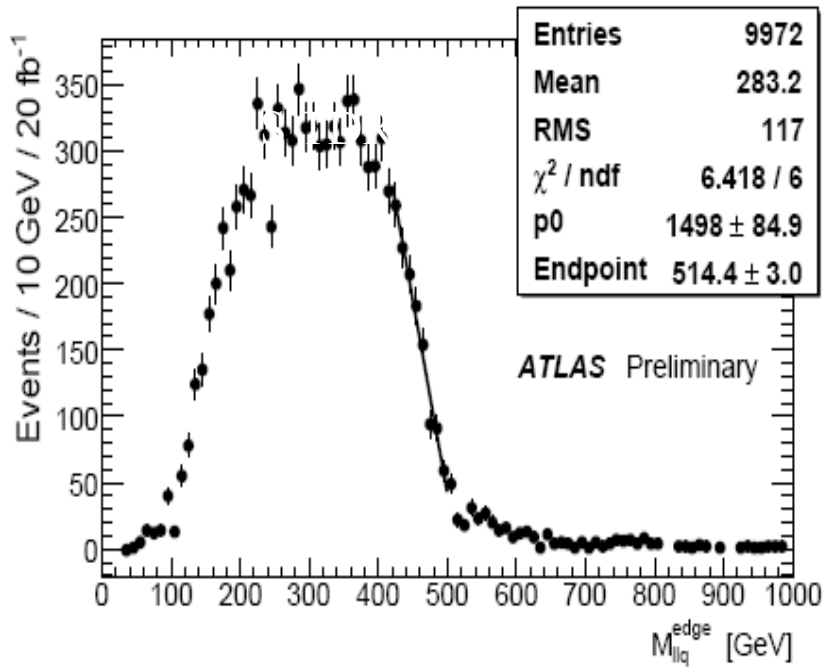
# Lepton+jets combination



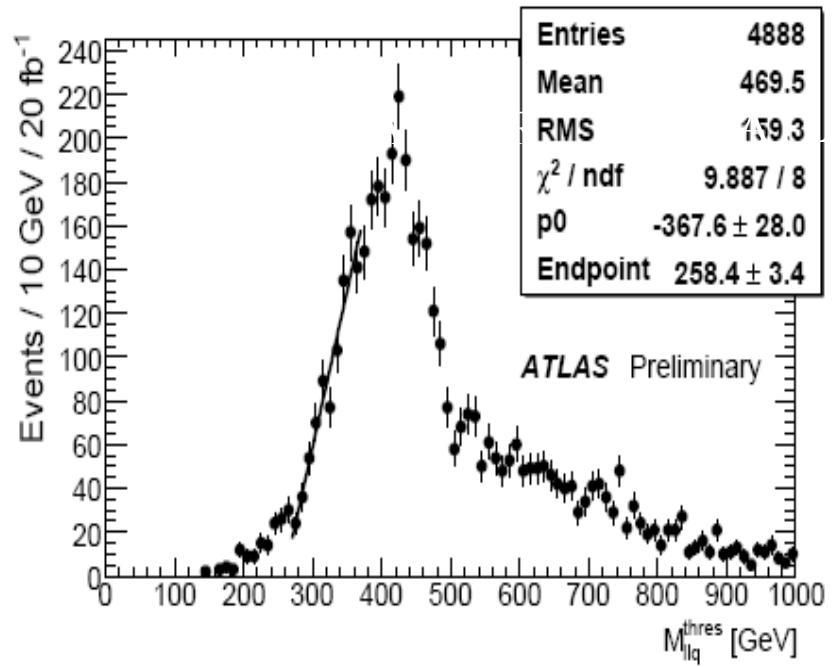
Formulas in Allanach et al., hep-ph/0007009

The invariant mass of each combination has a minimum or a maximum which provides one constraint on the masses of  $\tilde{\chi}^0_1$ ,  $\tilde{\chi}^0_2$ ,  $\tilde{l}$ ,  $\tilde{q}$

***llq edge***



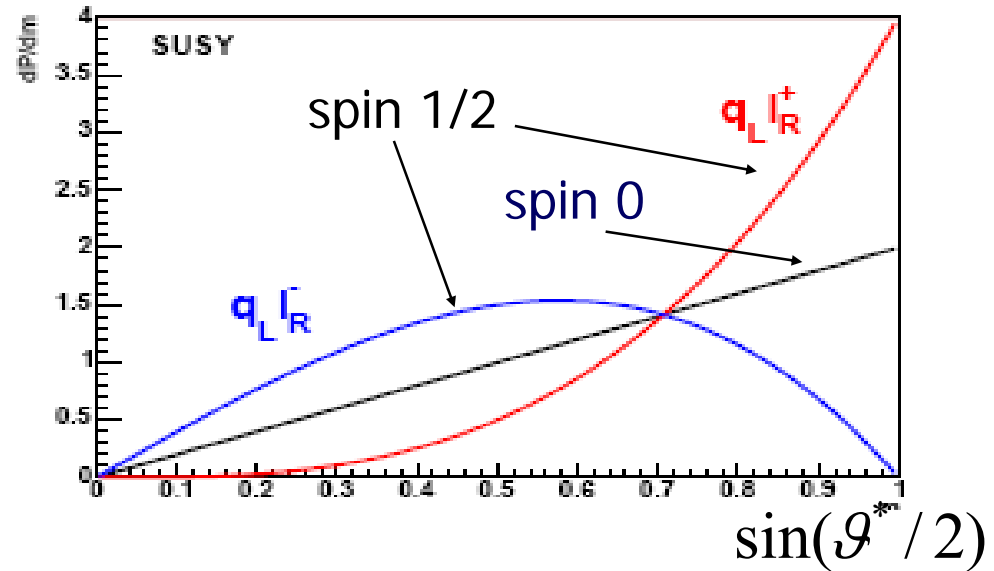
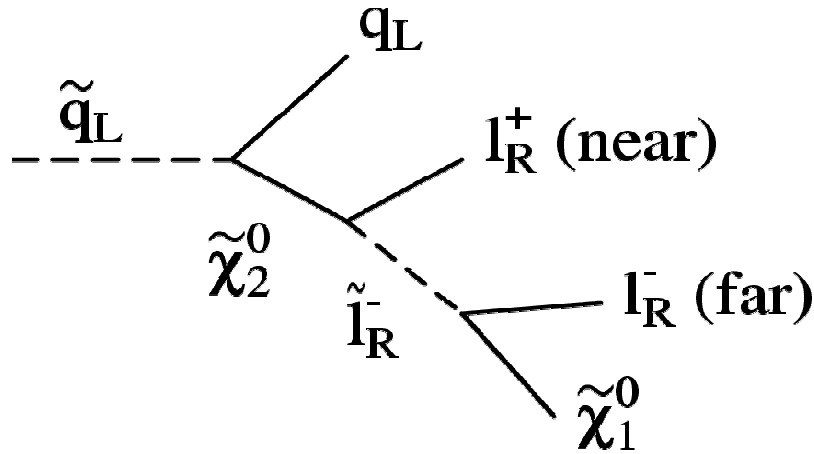
***llq threshold***



# Measurement of neutralino spin (1)



Important to measure the spin of new particles: it's the fundamental check to ensure that what we have discovered is SUSY!!



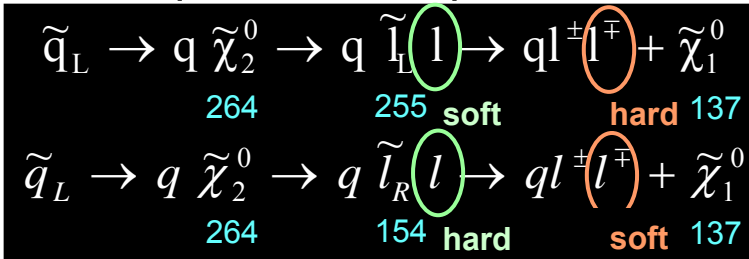
The charge asymmetry is **diluted** because:

1. Usually it is not possible to discriminate the *near* and *far* leptons: we sum  $m(q l^{\text{far}})$  and  $m(q l^{\text{near}})$  invariant masses
2. The charge conjugated cascade decay (from the anti-squark) gives the opposite asymmetry. However, cancelation is not exact because at LHC a larger number of squarks than anti-squarks is produced (pp collider)

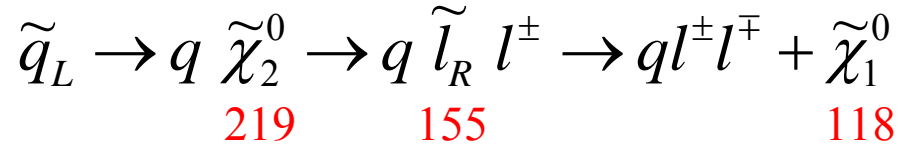
# Measurement of neutralino spin (2)



SU1 point: 7.8 pb x 1.6%  
 Ratio squarks/anti-squarks ~3.5

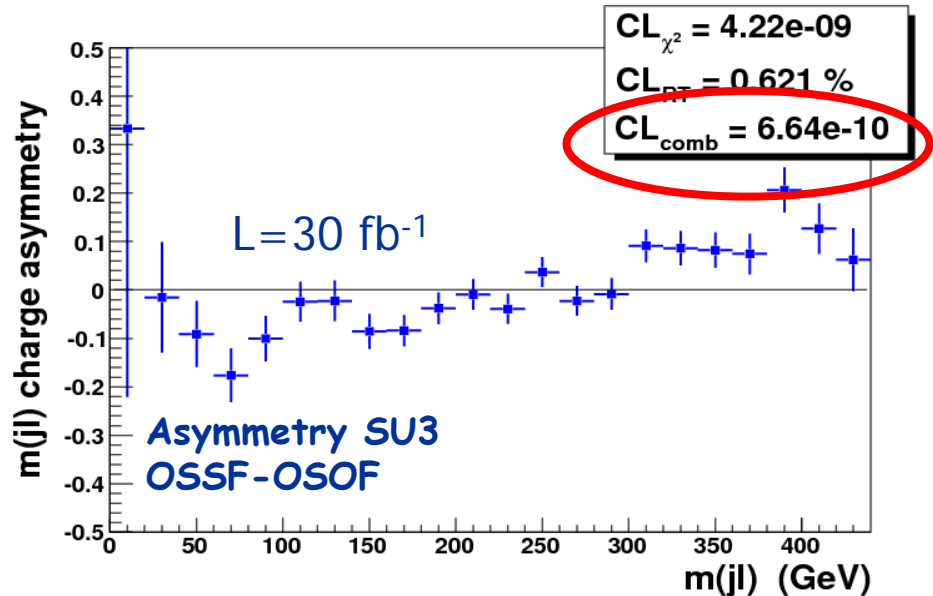
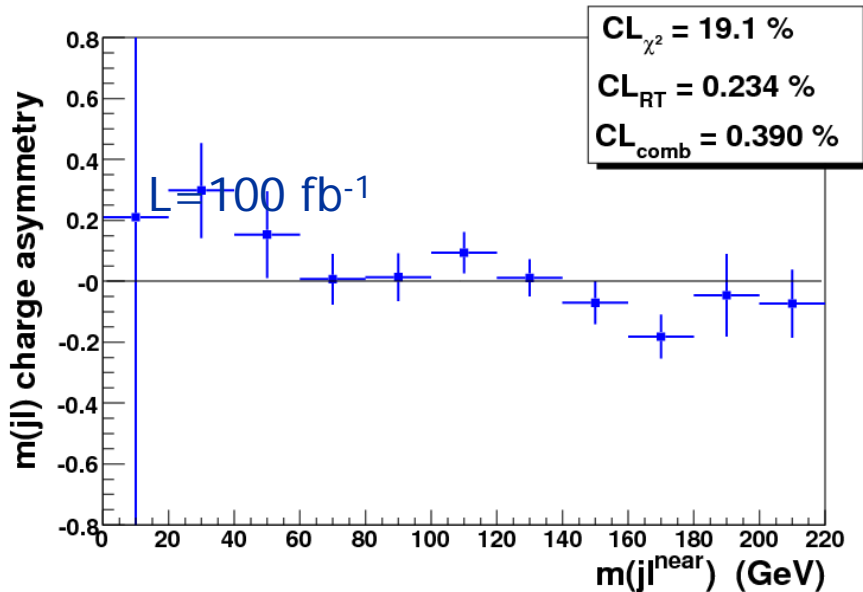


SU3 point: 19.3 pb x 3.8%  
 Ratio squarks/anti-squarks ~3



- Cuts on missing energy and jet pt to reject SM background
- 2 Opposite Sign, Same Flavour (OSSF) electrons or muons.
- Subtract background from independent decay chains with the combination  $\mu^+ \mu^- + e^+ e^- - \mu^\pm e^\mp$

In **SU3** point, **5-10 fb<sup>-1</sup>** are **already enough** to exclude charge symmetry



## Other SUSY scenarios



Across the MSSM, there is a rich variation in the SUSY phenomenology. The signatures expected at the LHC can be very different from the “mainstream” scenario discussed so far.

- **GMSB**: the lightest SUSY particle is the gravitino. The next-to-lightest particle (NLSP) decay only through gravitational interactions and may live longer than the time-of-flight across the detector.
- **Split SUSY**: scalars are much heavier than the electroweak scale. The gluino decays through virtual squarks, and may live longer than the time of flight across the detector.
- **R-parity violation**: the neutralino decays. Less missing energy and more jets or other particles.
- **Light stop models**: a scalar top with 120-150 GeV mass is still allowed.

Each scenario is covered by dedicated search strategies. I will discuss the GMSB scenario here...



- **G**auge **M**ediated **S**upersymmetry **B**reaking.  
Models for **SUSY breaking**, alternative to **mSUGRA**
- SUSY breaking transmitted from Hidden sector to visible sector **via gauge interactions** (“messengers”)
- Why interesting?
  - more **natural suppression of FCNC**
  - not huge  $\sigma$  but **clear signature** to claim early discovery or exclusion
    - $\sigma \sim 0.1 \div 1$  pb (model dependent)

Par.	Description
$\Lambda$	SUSY breaking scale
$M_m$	Messenger mass scale
$\tan\beta$	Ratio of Higgs vev
$N_m$	Number of SU(5) messenger multiplets
$\text{sign}(\mu)$	$\mu$ from Higgs sector
$C_{\text{grav}}$	Sets NLSP lifetime

- LSP is the **Gravitino** ( $m \leq \text{keV}$ )
  - light, stable and weakly interacting
  - possible candidate for Dark Matter

Present limits: Tevatron,  $\Lambda > 80$  TeV,  
 $m(\text{neutralino}, \text{chargino}) > 108, 195$  GeV

# GMSB Model (2)



Phenomenology depends on nature and lifetime of the second lightest state (NLSP):

	$\tilde{\tau}$ or $\tilde{l}$ is NLSP	$\tilde{\chi}_1$ is NLSP
$c\tau \gg L$	Like an heavy $\mu$	Like mSUGRA
$c\tau \approx L$	NLSP decays in the detector, lifetimes measurements.	
$c\tau \ll L$	Decays into 2 $\tau$	Decays into 2 $\gamma$

- **$\tau$  trigger and reconstruction** in early data not trivial
- **Decay into 2 $\gamma$  promising** (good ECAL performance early enough?)
- **Lifetime measurements**: need to understand **vertexing** in early data
  - For longer lifetimes, need to understand **background**:
    - Hard radiation from high- $p_T$  cosmic muons
    - Delayed hadronic showers ( $K_L^0$  and neutrons)

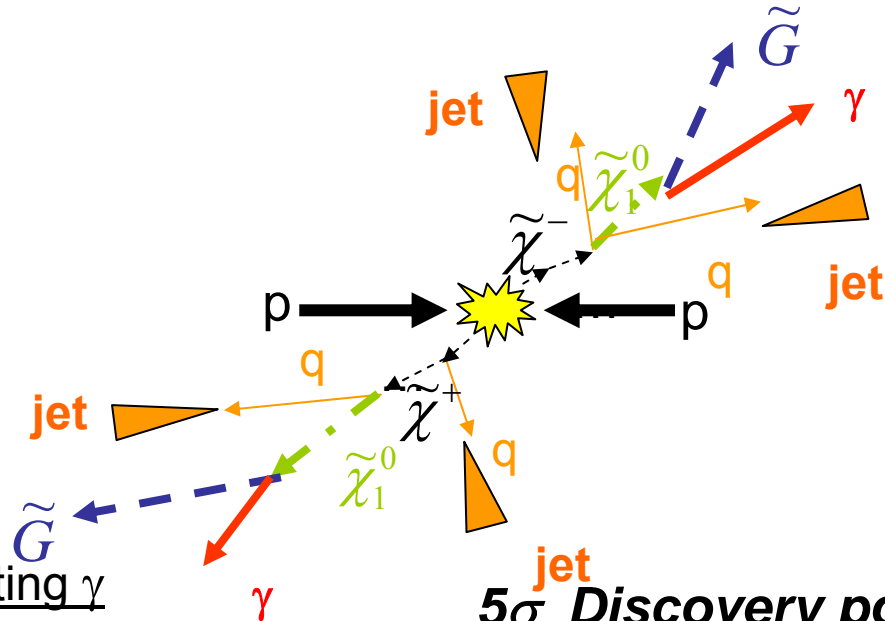




# GMSB Model: Performances (1)

MET  
+ photons

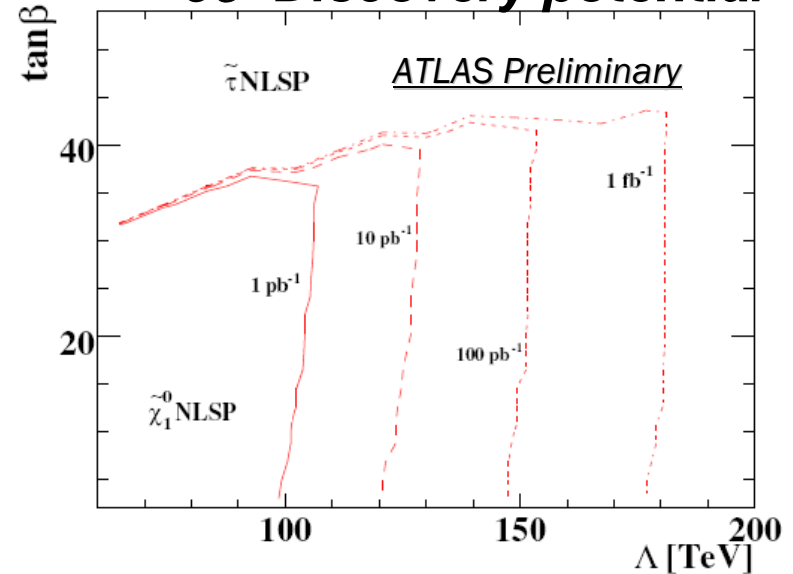
- If NLSP is **neutralino**  
⇒  $2\gamma$  in event
- Selection
  - $\gamma$ , isolation,  $P_T > 80$  GeV
  - High MET,  $N_{\text{jets}} > 3$
- **Main backgrounds**
  - $\gamma$ +jets
  - W+jets
- If  $\text{lifetime}(\chi) \neq 0 \Rightarrow$  non-pointing  $\gamma$   
⇒ possible to extract **lifetime**



**MET tails critical** for early discoveries

- Trigger efficiency (combining jet, MET and photon triggers) seems not a problem at  $10^{33}$  luminosity menus.
- Possible bias in lifetime measure from identification and reconstruction cuts for photons.

## $5\sigma$ Discovery potential

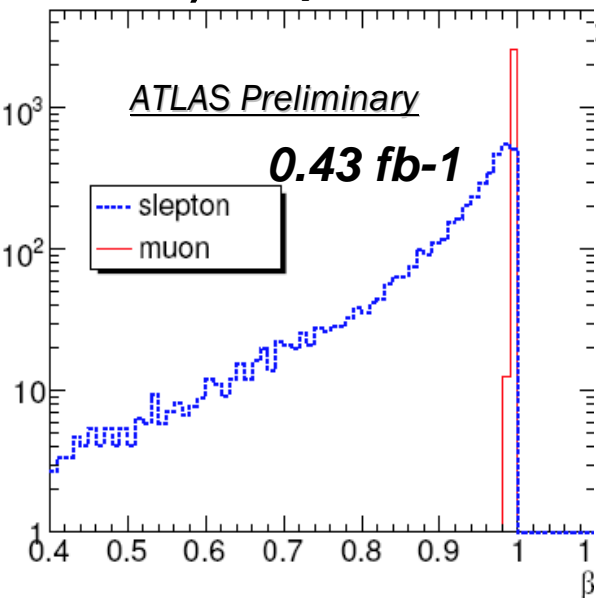


# GMSB Model: Performances (2)

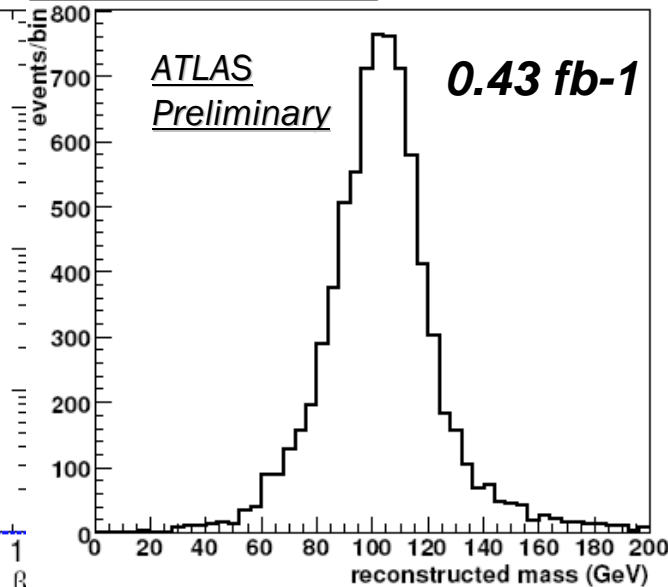


- Heavy slow “stable” leptons can be tagged with Time-Of-Flight measurements in muon drift tubes.
- Large calorimetric  $E_{\text{MISS}}^T$  due to quasi-stable leptons, like in mSUGRA.
- Timing/trigging issues most critical (association to the correct BCID problematic if  $\beta < 0.7$ , recoverable with MDT but specific algorithm for long-lived heavy particles will be useful).

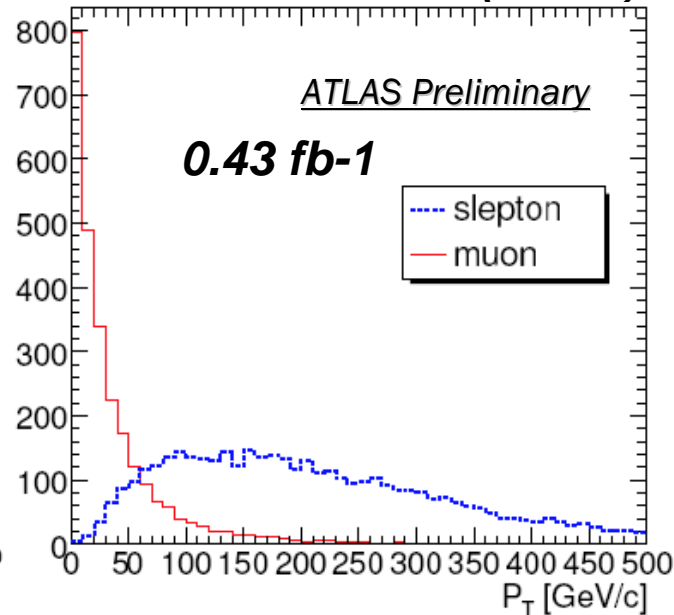
$\beta$  slepton



slepton mass peak



Transv.momentum (Gev/c)



$$\Lambda = 30\text{TeV}, M_m = 250\text{TeV}, N_m = 3, \tan \beta = 5, \text{sign}(\mu) = +, C_{grav} = 4$$

$$m(q, g) \approx 700\text{GeV} \quad m(\chi) \approx 114\text{ GeV} \quad m(l) \approx 102\text{ GeV} \quad \sigma = 23\text{ pb}$$

# Conclusions



- A brief review of the search strategies for SUSY in ATLAS has been presented;
  - New discoveries possible with early LHC data ( $O(100)\text{pb}^{-1}$ )
- Accurate knowledge of **SM physics** and of **detector performance** needed for any new discovery
  - **First data** taking period devoted to understanding of detector
- Any claim of new physics requires check of **trigger refinements** and data-driven estimates of **syst./background**
  - First, focus on less systematic-affected analyses (e.g. striking signatures and resonances)
- Larger statistics needed for full scan over SUSY parameters space and **discrimination** between different models.

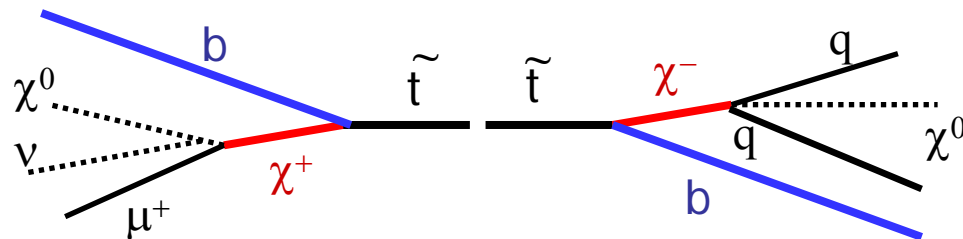


# BACKUP SLIDES



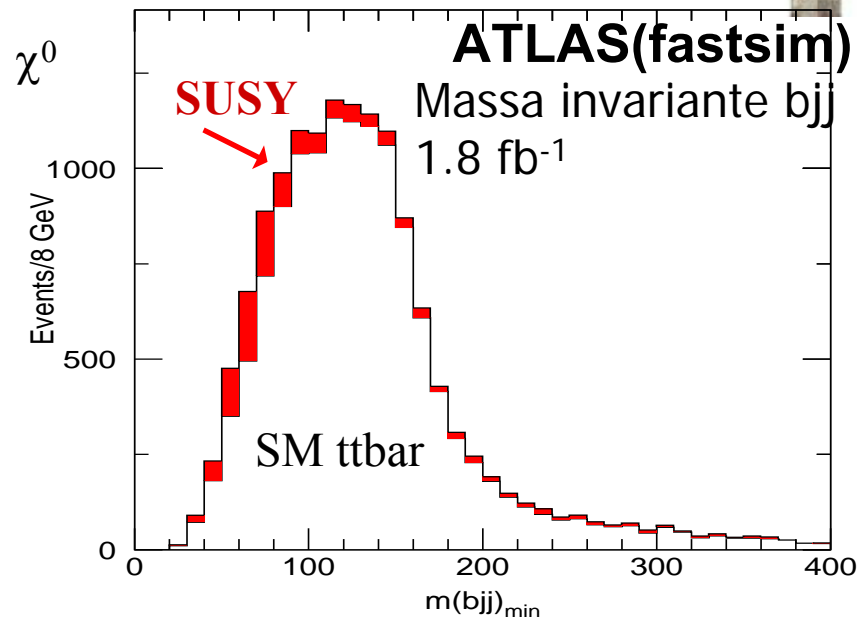
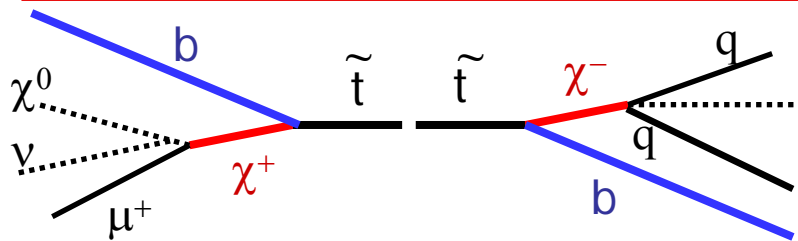
## Light stop scenario

- Direct limits allows the scalar top to be lighter than top
- There are models which explain baryogenesis (the generation of matter-antimatter asymmetry of the Universe) and Dark Matter at once using SUSY
  - Give up SUGRA-like unification of SUSY masses
  - Require a very light stop
  - ... and of course CP violation
- Consider direct production stop pairs:



- Looks a lot like top pair production
  - Cross section is comparable (400 pb for a 140 GeV stop)
  - Same final state, but “wrong” invariant mass combination (no W, top peak)
  - Still two unobserved neutralinos: no mass peak!
  - Softer leptons, jets and missing energy than in  $t\bar{t}$
  - Biggest problem is  $t\bar{t}$  background

# Search for light stop (I)



## Event selection:

4 jets with  $p_T > 25$  GeV

$E_T^{\text{miss}} > 20$  GeV

1 elec. or muon with  $p_T > 20$  GeV

$M(jj) < 60$  GeV (veto on  $W \rightarrow jj$ )

**Distribution of interest:**  $M(jjb)$ ,  $M(lb)$

The signal is “visible” on top of the SM background if we assume we know (from Montecarlo predictions) how many SM events (and the shape of distribution) pass event selection on average.

Since we may not trust the MC prediction to this level of accuracy, we developed a technique to estimate the shape of the SM contribution to the distribution.

Once we know the shape, we can fix the normalization of the background using the events at large invariant mass, where no SUSY contribution is expected.

# Search for light stop (II)

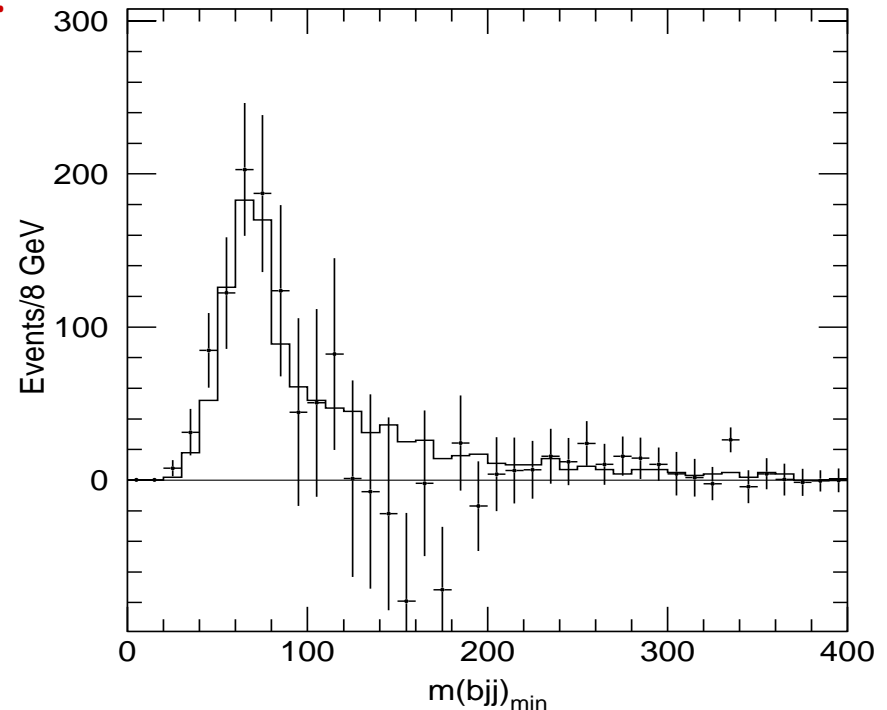


**We developed a technique to estimate the  $t\bar{t}b\bar{r}$  bckg from data:**

**Control sample 1:** tight extra cuts on hadronic side [ $M(jjb) = m(\text{top})$ ,  $M(jj)=M(W)$ ] select  $t\bar{t}b\bar{r}$   
- Used to measure shape of  $M(bl)$  in  $t\bar{t}b\bar{r}$  events

**Control sample 2:** tight cuts on leptonic side [ $M(lb,xEt) = M(\text{top})$ ] to select  $t\bar{t}b\bar{r}$   
-Used to measure shape of  $M(bjj)$  in  $t\bar{t}b\bar{r}$  events

**Signal visible after background subtraction with  $\sim 1 \text{ fb}^{-1}$**



**Solid line:** SUSY events among those passing event selection

**Points:** Measured distribution, after subtracting the SM contribution estimated with control samples

**It works!**



## *mSUGRA models*

- **A random choice of the 105 MSSM parameters violates limits** from B/D/K physics, electric dipole moments, FCNC, ...
- **Need some assumption** on the structure **of SUSY breaking lagrangian**. In **mSUGRA** (5 free parameters, most studied by ATLAS and CMS):
  - **Conserved R-parity**
  - **Common mass**  $m_0$  for susy scalars,  $m_{1/2}$  for fermions (at GUT scale).
  - **Common value**  $A_0$  for the trilinear coupling of the s-fermions with the 2 Higgs doublets.

**Then 5 free parameters:**  $m_0, m_{1/2}, A_0, \tan \beta, \text{sgn } \mu$

Further constraints if it is required that the Big Bang has produced the right amount of stable neutralinos to explain observed Dark Matter density

May be too constrained. Experiments at colliders are interested mostly in identify signatures to develop and study search strategies





# *Supersymmetry: what is?*

## Supersymmetry (SUSY) in a nutshell

### *Standard particles*

Quarks, leptons, neutrinos (spin 1/2)

W, Z, gluino (spin-1)

Higgs (spin-0)

### *Superpartners*

Squarks, sleptons, sneutrinos (spin-0)

Wino, zino, gluino (spin 1/2)

Higgsino (spin 1/2)

At least two Higgs doublets are needed → **five Higgs bosons**

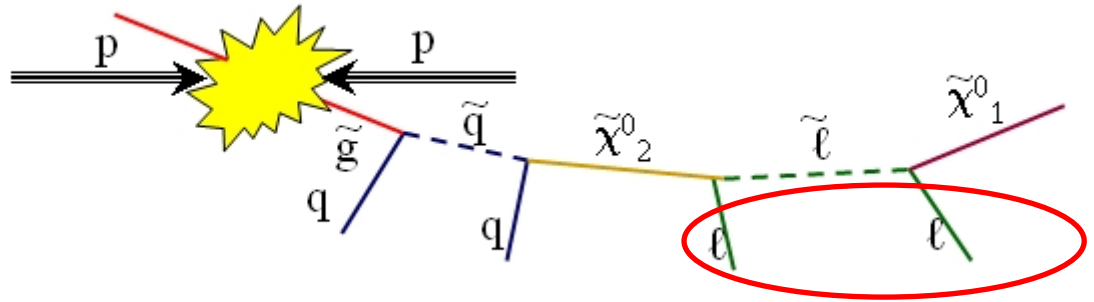
Wino, Zino, Higgsino mix → 4 charged (chargino) and 4 neutral (neutralino) states

SUSY particles not observed yet → must be heavy → **symmetry is broken**

It is possible to put directly SUSY mass terms in the lagrangian. This gives about **100 free parameters** with the minimal field content above (MSSM model)

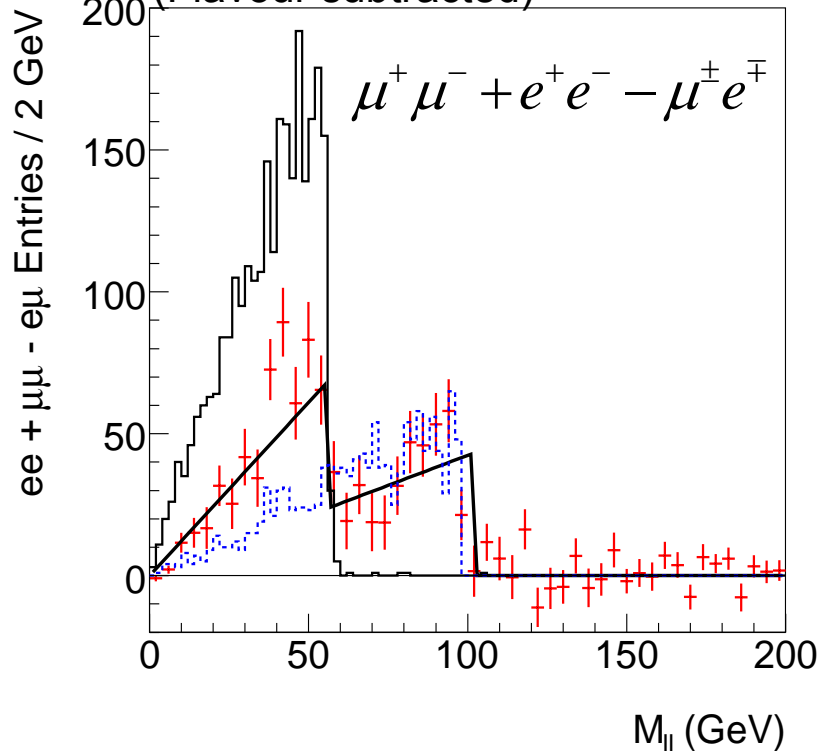
**Constrained models** (with assumptions on the structure of SUSY breaking) have only a few parameters – but assumptions may be wrong.

# Kinematical structures



Same-Flavour Other-sign  
Lepton combinations.

(Flavour subtracted)



The invariant mass of the two leptons has a kinematical endpoint which measures:

$$M_{ll}^{\max} = M(\tilde{\chi}_2^0) \sqrt{1 - \frac{M^2(\tilde{l}_R)}{M^2(\tilde{\chi}_2^0)}} \sqrt{1 - \frac{M^2(\tilde{\chi}_1^0)}{M^2(\tilde{l}_R)}}$$

It may be possible to observe two edges, if both decays are open:

$$\begin{aligned} \tilde{\chi}_2^0 &\rightarrow \tilde{l} \tilde{l}_L \rightarrow ll \tilde{\chi}_1^0 \\ \tilde{\chi}_2^0 &\rightarrow \tilde{l} \tilde{l}_R \rightarrow ll \tilde{\chi}_1^0 \end{aligned}$$

The SM and SUSY combinatorial backgrounds have two leptons from independent decay chains. The background cancel in the flavour subtraction



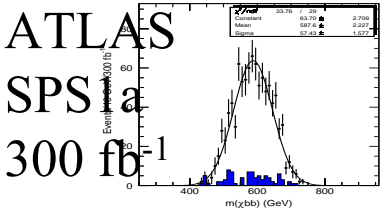
# Getting mass peaks

- The 4-momentum of the  $\chi^0_2$  can be reconstructed from the approximate relation

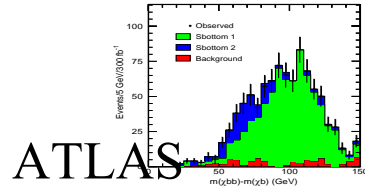
$$p(\chi^0_2) = ( 1 - m(\chi^0_1)/m(l\bar{l}) ) p_{ll}$$

valid when  $m(l\bar{l})$  near the edge.

- The  $\chi^0_2$  can be combined with b-jets to reconstruct the gluino and sbottom mass peaks from  $\tilde{g} \rightarrow b\bar{b} \rightarrow bb\chi^0_2$



$m(\chi_{bb})$  (GeV)



ATLAS  
SPS1a  
300 fb<sup>-1</sup>

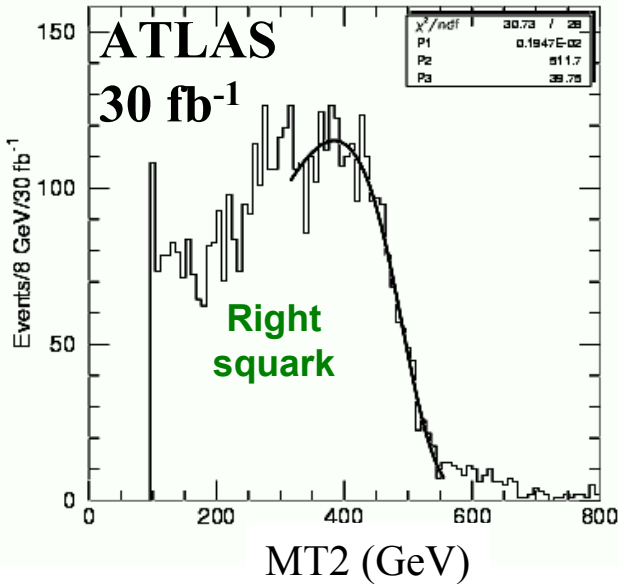
$m(\chi_{bb}) - m(\chi_b)$  (GeV)

SPS1a, 300 fb<sup>-1</sup>, stat. errors only:  
 $m(\tilde{g}) - 0.99m(\chi^0_1) = (500.0 \pm 6.4) \text{ GeV}$   
 $m(\tilde{g}) - m(\tilde{b}_1) = (103.3 \pm 1.8) \text{ GeV}$   
 $m(\tilde{g}) - m(\tilde{b}_2) = (70.6 \pm 2.6) \text{ GeV}$

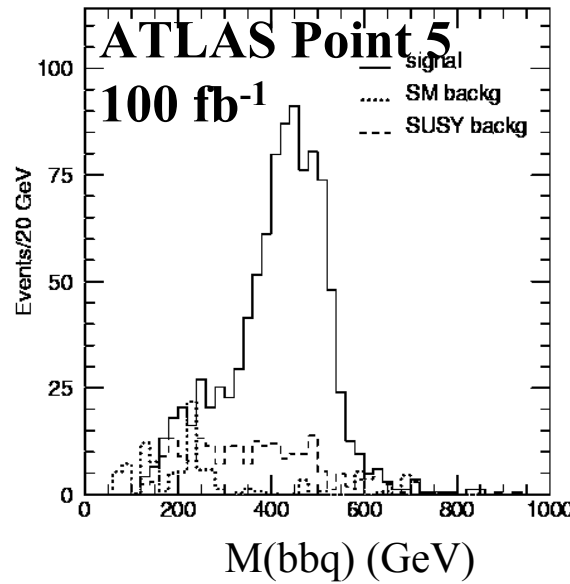


# Other mass measurements

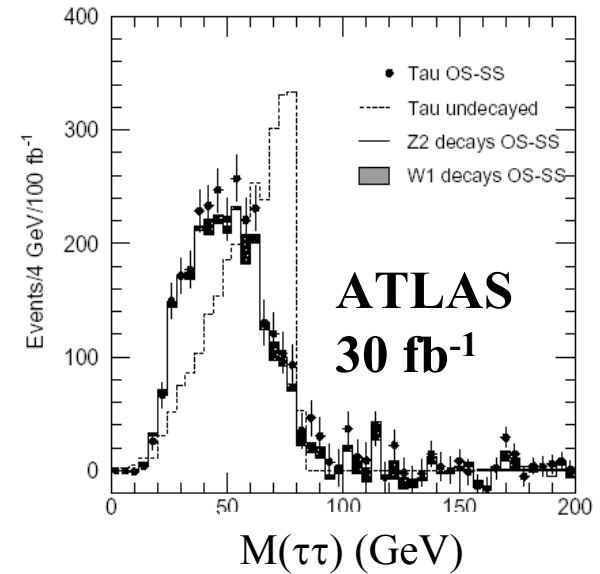
$$q_R \rightarrow \chi^0_1 q$$



$$q_L \rightarrow \chi^0_2 q \rightarrow \chi^0_1 h q \rightarrow \chi^0_1 b b q$$



$$\chi^0_2 \rightarrow \tilde{\tau} \tau \rightarrow \chi^0_1 \tau \tau$$



2 hard jets and lots of  $E_T^{\text{miss}}$ .

Reconstruct with

$$M_{T2}^2 = \min_{p_1+p_2=p_T} \left[ \max \left\{ m_T^2(p_T^{\ell_1}, p_1), m_T^2(p_T^{\ell_2}, p_2) \right\} \right]$$

$$m(\tilde{q}_R) - m(\chi^0_1) = (424.2 \pm 10.9) \text{ GeV}$$

Also works for sleptons.

Two body decay of  $\chi^0_2$  to higgs and  $\chi^0_1$ .

Reconstruct higgs mass (2 b-jets) and combine with hard jet.

Get additional mass

constraint

Tau decay dominates neutralino BR at large  $\tan\beta$ .

No sharp edge because of  $\nu$ , but end-point can still be measured.

# Supersymmetry: why?



Supersymmetry can solve several problems of the Standard Model at once

## Hierarchy problem:

- Fermions and bosons contribute with opposite sign to the Higgs mass
- $\delta m_H \sim m_{\text{SUSY}}$  [SUSY mass scale]
- Hierarchy ok if SUSY masses near the Higgs scale (**accessible to a TeV-scale collider**)

True also for other SM extensions addressing hierarchy.

**The TeV-scale new physics and the Higgs are the main motivations for the Large Hadron Collider**

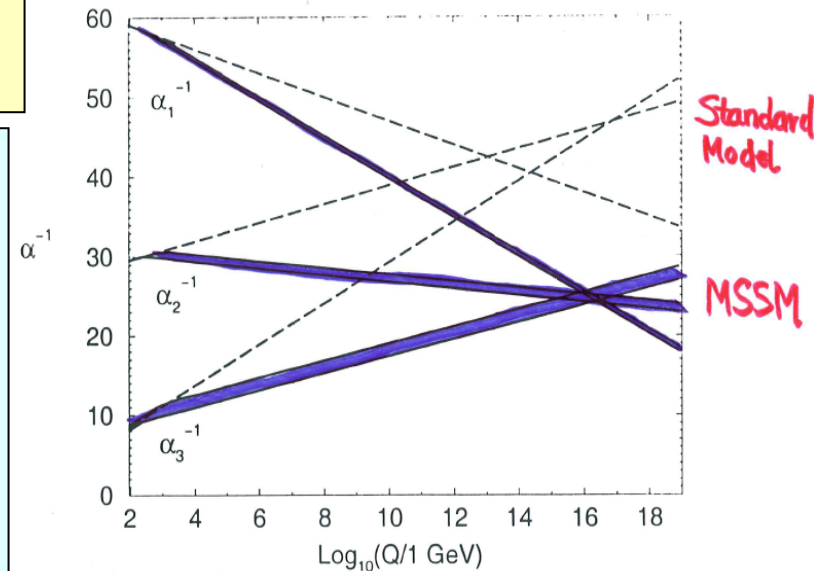
## Dark Matter

Need a conserved quantum number to avoid proton decay:  $R = +1$  for SM particles,  $R = -1$  for SUSY particles. Consequences:

- **SUSY particles are produced in pairs**
- **The lightest SUSY particle is stable.** If weakly interacting, it's a good candidate for Dark Matter

## Unification of forces:

Better convergence of interaction strength as a function of energy



# HT2 method

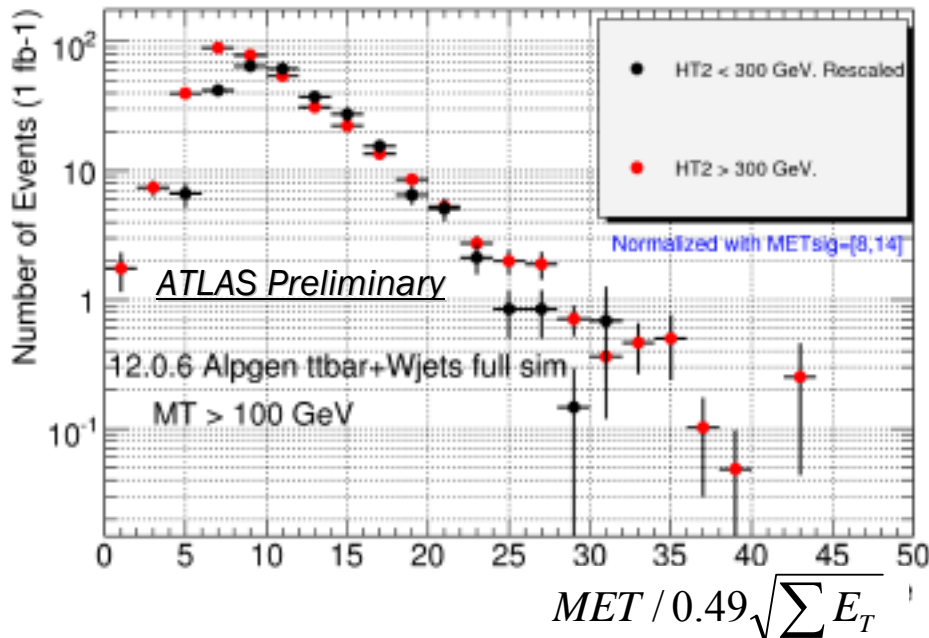
An other variable which has small correlation with MET is

$$HT2 \equiv \sum(\text{pt jets } 2,3,4) + \sum(\text{pt } e, \mu)$$

- leading jet is not included in order to avoid correlation with MET
- use MET significance rather than MET to reduce correlation

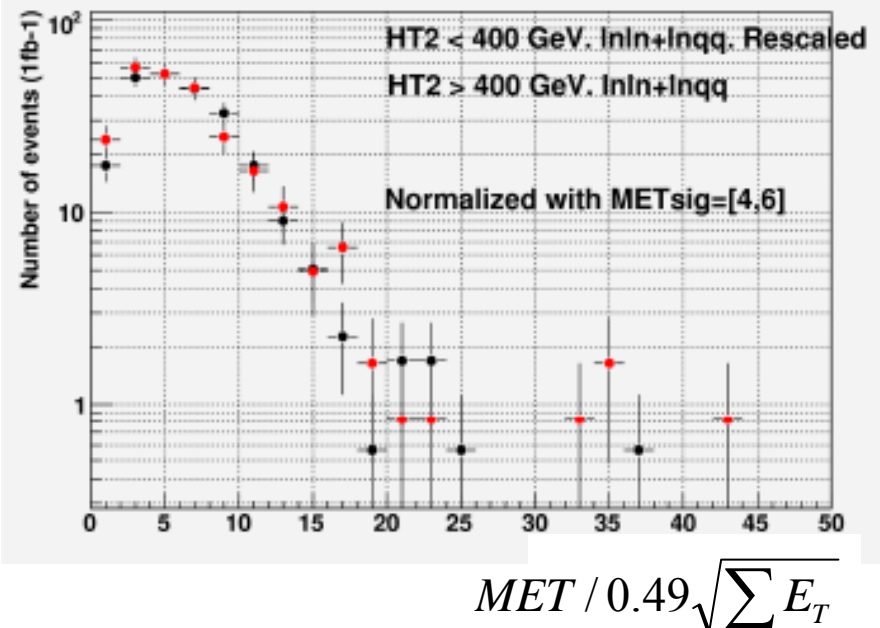
## one lepton mode

*t*t̄bar (ℓνℓν and ℓνqq), W(ev)+jets, W(μν)+jets



## also works for OS di-lepton mode

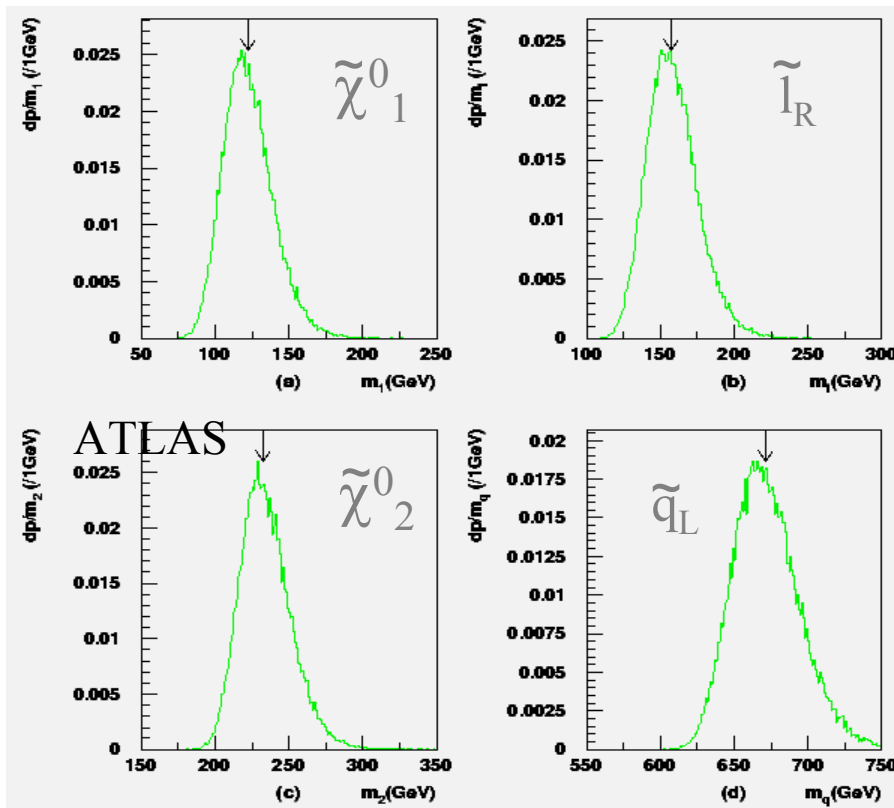
*ATLAS Preliminary*



# Getting SUSY particle masses



- Combine measurements from edges of different jet/lepton combinations to obtain 'model-independent' mass measurements.
- LSP mass uncertainty large, all other masses strongly correlated with it. A future Linear Collider measurement of  $\chi^0_1$  mass would improve the precision on all masses.



masses (GeV)	LHCC5	SPS1a
$m(\tilde{\chi}^0_1)$	122	96
$m(\tilde{l}_R)$	157	143
$m(\tilde{\chi}^0_2)$	233	177
$m(\tilde{q}_L)$	687-690	537-543

Sparticle	Expected precision (100 fb <sup>-1</sup> )
$\tilde{q}_L$	$\pm 3\%$
$\tilde{\chi}^0_2$	$\pm 6\%$
$\tilde{l}_R$	$\pm 9\%$
$\tilde{\chi}^0_1$	$\pm 12\%$

# From masses to model parameters

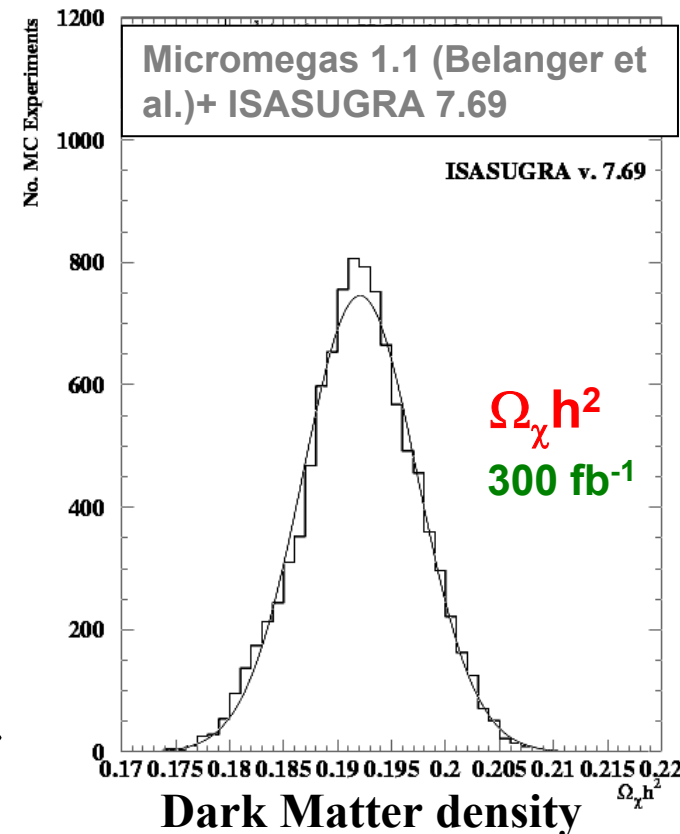


From a given set of measurements one scans the **parameter space** and finds the points compatible with data. These points are fed to relic density calculators to get constraints on **neutralino dark matter abundance**

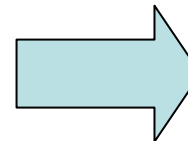
Variable	Value (GeV)	Errors		
		Stat. (GeV)	Scale (GeV)	Total
$m_{\ell\ell}^{max}$	77.07	0.03	0.08	0.08
$m_{\ell\ell q}^{max}$	428.5	1.4	4.3	4.5
$m_{\ell q}^{high}$	300.3	0.9	3.0	3.1
$m_{\ell q}^{high}$	378.0	1.0	3.8	3.9
$m_{\ell\ell q}^{min}$	201.9	1.6	2.0	2.6
$m_{\ell\ell b}^{min}$	183.1	3.6	1.8	4.1
$m(\ell_L) - m(\tilde{\chi}_1^0)$	106.1	1.6	0.1	1.6
$m_{\ell\ell}^{max}(\tilde{\chi}_4^0)$	280.9	2.3	0.3	2.3
$m_{\tau\tau}^{max}$	80.6	5.0	0.8	5.1
$m(\tilde{g}) - 0.99 \times m(\tilde{\chi}_1^0)$	500.0	2.3	6.0	6.4
$m(\tilde{q}_R) - m(\tilde{\chi}_1^0)$	424.2	10.0	4.2	10.9
$m(\tilde{g}) - m(\tilde{b}_1)$	103.3	1.5	1.0	1.8
$m(\tilde{g}) - m(\tilde{b}_2)$	70.6	2.5	0.7	2.6

$$\Omega_\chi h^2 = 0.1921 \pm 0.0053$$

$$\log_{10}(\sigma_{\chi p}/\text{pb}) = -8.17 \pm 0.04$$



Parameter	Expected precision (300 fb <sup>-1</sup> )
$m_0$	$\pm 2\%$
$m_{1/2}$	$\pm 0.6\%$
$\tan(\beta)$	$\pm 9\%$
$A_0$	$\pm 16\%$





## *How much data will we need?*



Statistical reach with  $100 \text{ pb}^{-1}$  is in the TeV region, well beyond Tevatron limits ( $\sim 400 \text{ GeV}$ ) BUT

- **only in a few cases SUSY has distinctive kinematical features**
- **main selection tool at both trigger and analysis level is to select event with large missing  $E_t$ , difficult to muster experimentally**

**More luminosity (for control samples) and/or time may be needed to understand backgrounds**

**Let's go back to detector commissioning and SM background studies...**